

Aircraft Observations of Thin Cirrus Clouds near the Tropical Tropopause

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Abstract. This work describes aircraft-based lidar observations of thin cirrus clouds at the tropical tropopause in the central Pacific obtained during the Tropical Ozone Transport Experiment/Vortex Ozone Transport Experiment (TOTE/VOTE) in December 1995 and February 1996. Thin cirrus clouds were found at the tropopause on each of the four flights which penetrated within 15 degrees of the equator at 200-210 east longitude. The altitudes of these clouds exceeded 18 km at times. The cirrus observations could be divided into two basic types – thin quasi-laminar wisps and thicker, more textured structures. Based on trajectory analyses and temperature histories, these two types were usually formed respectively by: (1) *in situ* cooling on both a synoptic scale and mesoscale; and (2) recent (a few days) outflow from convection. There is evidence from one case that the thicker clouds can also be formed by *in situ* cooling. The actual presence or absence of thin cirrus clouds was also consistent the temperature and convective histories derived from back trajectory calculations. Notably, at any given time, only a relatively small portion (at most 25%) of the west central tropical Pacific has been influenced by convection within the previous 10 days. The structures of some of the thin cirrus clouds formed *in situ* strongly resembled long-wavelength (500 - 1000 km) gravity waves observed nearly simultaneously by the ER-2 on one of the flights. Comparison with *in situ* water vapor profiles made by the NASA ER-2 aircraft provide some observational support for the fact that thin cirrus clouds may play an important role in dehydrating tropospheric air as it enters the stratosphere.

1. Introduction

Recent work by *Jensen et al.*, [1996] has shown that subvisible and horizontally extensive subvisible cirrus clouds (SVC) near the tropical tropopause can govern the water vapor input into the lower stratosphere by simultaneously removing water vapor by sedimentation and lofting the affected air parcels through enhanced IR radiative heating of cloud particles. The large horizontal extent of SVC's (several 100 km, as shown in McFarquhar et al., [1999]) and substantial frequency over large parts of the tropical atmosphere [*Wang et al.*, 1996] implies a role in: (1) the thermal balance of the tropical tropopause region; and (2) the earth's overall radiation balance. As shown by *Rosenfield et al.*, [1997], cirrus formation may raise the average temperature of the tropical tropopause by 1-2K. Since such a temperature increase implies an increase in the ice saturation mixing ratio, SVC formation at the tropical tropopause is raising the water vapor input into the lower stratosphere above what it would otherwise be. Recent radiative forcing calculations (Jensen, personal communication) show that an atmosphere with a "wet" tropopause (80% relative humidity) has 1 wattm^{-2} less outgoing longwave radiation than a "dry" tropopause (50% relative humidity), a significant number for net radiative forcing in global climate studies. Clouds, even thin ones, would substantially enhance this difference.

The potential importance of these clouds to the stratospheric water budget depends on how they are formed. There are two possibilities: excess ice left over from a convective outflow and *in situ* formation. If they are excess ice left over from a

convective outflow (i.e., ice formed in the convective system and passed into the anvil outflow that has failed to fall out because of small particle size), then their role depends on the position of that convective outflow with respect to the tropopause. If they are below the tropopause, they will be radiatively heated, but their temperature will still decrease as the cloud rises in an environment where temperature decreases with altitude [Ackerman *et al.*, 1988]. In this case, the convective outflow ice crystals will continue to grow and presumably fall out, and the clouds will act both to advect moisture upward and simultaneously remove it [Sherwood, 1999]. Given that *in situ* ice nucleation requires some degree of supersaturation [Heymsfield *et al.*, 1998], the presence of "seed" ice from convection may actually enhance dehydration by allowing ice particles to grow at saturated rather than supersaturated water vapor values. Above the tropopause, the reverse is true. Radiative heating of the convective outflow will simply evaporate the leftover ice crystals and act to hydrate the lower stratosphere. If they are formed *in situ* (due to temperature changes induced by wave activity or large scale ascent in the tropopause region), they can be viewed as a mechanism that simultaneously lifts air upward and strips it of water vapor [Jensen *et al.*, 1996]. It should be noted that convective outflow cirrus can also be lifted and cooled by wave activity, leading to net dehydration. The observations documented in this paper indicate the presence of both convective outflow cirrus and cirrus formed *in situ*. These observations also show that short-term (10 days or less) trajectories near the tropopause based on standard analyses are quite credible even in the data-sparse tropical regions.

2. Experimental and Meteorological Background

The Tropical Ozone Transport Experiment/Vortex Ozone Transport Experiment (TOTE/VOTE), was actually designed to examine filaments of stratospheric tracers and assess their role in large scale lateral mixing in the tropics and polar regions. To examine these filaments, NASA's DC-8 aircraft was equipped with a number of *in situ* and remote instruments, including a UV Differential Absorption Lidar (UVDIAL – *Browell et al.*, [1983] and *Browell et al.*, [1998] measuring aerosol backscatter, aerosol depolarization, and ozone, and a Microwave Temperature Profiler (MTP – *Denning et al.*, [1989] and *Gary*, [1989]), measuring the temperature profile between about 5 km and 25 km altitude along the aircraft flight track. The experimental protocol called for the aircraft to fly at maximum cruise altitude (about 11 km) at night (to maximize lidar signal to noise ratios) and make measurements in a vertical profile between the aircraft and maximum instrument altitude (about 25 km for UVDIAL ozone and MTP temperature). Though an investigation of high altitude cirrus clouds was not an initial central goal, this experiment was unique in high altitude cirrus investigations in two ways. First, it provided, simultaneously, cloud information (the total scattering ratio and depolarization at 603 nm from the UVDIAL) and temperature information (MTP). Second, for one of the flights *in situ* data from a NASA ER-2 aircraft was available nearly simultaneously – that is the ER-2 flew along essentially the same track about 16 hours after the DC-8.

The tropical portions of TOTE/VOTE took place in two phases, December 13-20,

1995 and February 13-17, 1996. During these two periods, there were 6 local flights out of Hawaii: four were directed southward reaching the equator, one spanned the latitude range $13 - 28^{\circ}N$ with Hawaii as the midpoint, while one was directed northwestward of Hawaii. Figure 1a and Figure 1b show the flight tracks of the four equatorial flights superimposed on the average brightness temperature during the December, 1995 and February 1996 periods. Consistent with longer term climatologies (e.g., ISCCP, *Rossow and Schiffer*, [1991], the region south of Hawaii is relatively free of deep high clouds. The areas of colder average brightness temperatures coincide with the regions of tropical deep convection. For this period, these are, in decreasing order of overall convective activity : (1) a broad, region of extensive convection over Indonesia, northern Australia, and the eastern Indian ocean; (2) a line of convective activity in the southern hemisphere with a slight northwest-southeast tilt extending from Indonesia eastward across the international dateline towards American Samoa and south of Tahiti, also known as the Southern Pacific Convergent Zone (SPCZ); and (3) a narrow, fairly weak InterTropical Convergent Zone (ITCZ) extending eastward through the international dateline centered at about $7^{\circ}N$.

Figure 1a

Figure 1b

Also plotted in Figures 1a and 1b are the average wind fields and temperatures at 85 hPa (slightly above the tropopause level). Of note for this study is the monsoon anticyclone in the northern hemisphere extending from the equator to the subtropical jet edge near $20^{\circ}N$ and from about $90^{\circ}E$ to Hawaii. Outflow from region 1 convection flows (in part) northward toward the subtropical jet. It is then carried eastward by that jet and returns to the equator in a broad region of longitudes centered roughly

at the Hawaiian islands. Whereas the convective outflow and the subtropical jet west of the dateline are fairly steady, the return flow fluctuates substantially as amplifying waves travel along the subtropical jet, slow down as the jet weakens in the eastern Pacific, and "dig" equatorward. Along the equator, between Hawaii and the Indonesian convective region, the flow is westward, completing the anticyclone. A similar monsoon anticyclone is found in the southern hemisphere, centered roughly over the Australian continent. At Hawaiian longitudes, however, between the SPCZ and the equator, there is a weak northeastward flow which meets the southward branch of the northern anticyclone in a diffuent region near the equator. The temperature field shows that the coldest tropopause temperatures are centered at about $[12^{\circ}N, 150^{\circ}E]$, and do not, in fact, coincide with the regions of strongest convection. Thus, outflow from Indonesian convection may actually cool somewhat as it passes northward, then warm as it approaches the subtropical jet. In the return flow near Hawaiian longitudes, the air will cool substantially as it approaches the equator.

3. Observations

High altitude cirrus (HAC) were observed on all four southward flights, and very briefly on the Hawaii midpoint flight south of $15^{\circ}N$. The HAC were generally south of $15^{\circ}N$, extending as far north as $18^{\circ}N$ on the flight of 951215 (year, month and day). Here, we define thin cirrus as follows: (1) Total Scattering Ratio greater than 1.25; (2) 603 nm depolarization greater than .3; and (3) no evidence of high cloud (above the DC-8) based on the 6.5 and 10.5 micron channels from the geostationary meteorological

satellite imagery. Criterion (1) indicates the presence of a layer of particles distinct from the stratospheric aerosol (which has lower total scattering ratios) and criterion (2) indicates the degree of nonsphericity of those particles (with sufficiently high values indicating the presence of ice – *Browell et al.*, [1983]. Criterion (3) is a crude measure of optical depth – if it is met, whatever clouds are present are too thin to be detected by nadir sounders. Note here that the use of the term "subvisible cirrus" is avoided, since that term implies a maximum optical depth of .03 [*Sassen et al.*, 1989], a quantity that cannot be unambiguously determined from the available data. With few exceptions, the aircraft flew in nearly cloud free zones (essentially by design in order to allow the lidar instruments to make stratospheric observations).

Figures 2(a-h) show latitude-altitude cross-sections of the total scattering ratios observed by the DC-8 on the four equatorial Hawaii flights, with the southbound legs (a, c, e, and g) at left and the northbound legs (b, d, f, and h) at right. Also plotted are the temperatures from the Microwave Temperature Profiler [*Denning et al.*, 1989] and the potential temperatures from the Goddard GEOS-1 Assimilation Model [*Schubert and Rood*, 1993]. Note that in this and all subsequent figures, "altitude" refers to standard pressure altitude; this is equal to actual geometric altitude if the temperature profile were a US standard atmosphere temperature profile [*U.S. Standard Atmosphere Supplements* 1966]. Near the tropical tropopause, the actual geometric altitude is about .5 km higher than the standard pressure altitude. The corresponding nadir sounding images are in Figure 3 (a-d) (10.5 micron channel), which clearly show that all four flight tracks were well outside of significant convection.

Figures 2(a-h)

Figure 3 (a-d)

The aircraft essentially ascended in steps as fuel was burned, and these upward steps are indicated by the upward steps in the lower limit of the Total scattering ratio data. Three general features of the observations are to be noted. First, the HAC are clearly confined in altitude to the region within 2 km of the temperature minimum (presumably the tropopause), as shown by the MTP measurements. Second, the HAC seem to have two basic types. The first (HAC1) – as exemplified by the observations for the flight of 960213, the flight of 951220 north of $6^{\circ}N$, and the flight of 951215 north of $10^{\circ}N$ – are fairly uniform over large horizontal distances, have low total scattering ratios (less than 3), and small aspect ratios (their thickness to length ratio is .001 or less). They are also tilted, having a distinct slope of about .001 km/km for the HAC near $12^{\circ}N$ on 951220 and somewhat higher for 960213. Most of the tilted clouds sloped upward and poleward. However, on the northbound leg of the 960213 flight, there is an an opposite slope between $1 - 4^{\circ}N$ near 16 km. The second type – exemplified by the flight of 960217, the flight of 951220 south of $3^{\circ}S$, and the flight of 951215 south of $10^{\circ}N$ – has a blobbier, vertically thicker structure with no apparent slope (HAC2). The maximum total scattering ratios are larger than for HAC1, reaching 1000 for the HACs during the flight of 951215. The third general feature, not shown, is that all of the regions of enhanced total scattering ratio had significant depolarization, indicating that these were ice clouds with nonspherical particle shapes. Table 1 summarizes the locations and times that the two types of HAC were observed.

The observations also have some significant limitations, which make it difficult for this dataset to answer a question relevant to *Jensen et al's* mechanism – namely –

do these clouds occur above the tropopause? Unfortunately, because of the distance between the aircraft and the tropopause at these latitudes, the tropopause altitude yielded by the MTP is insufficiently precise to establish whether these clouds are unambiguously above the tropopause. Figure 2f, for example, clearly shows an upward step in the position of the minimum temperature that coincides with a step ascent by the aircraft. Since the accuracy of the tropopause altitude improves with the closeness of the MTP to the tropopause, we can say that the actual altitude of the minimum temperature point is probably higher than the MTP indicates. Based on this, all the thin cirrus observed on these four flights, with the possible exception of the example near 18km at $13.5^{\circ}N$ on 951215, are at or just below the tropopause.

There are two other aspects of the MTP measurements that are clearly reliable, and important for the present study. First, there are clearly significant mesoscale (horizontal scales of 100 km or less) fluctuations evident in the temperature contours in Figure 2. These fluctuations imply that an air parcel may experience peak to peak temperature changes on the mesoscale of 5K or more (Gary, personal communication). Second, in spite of its shortcomings in measuring the position and temperature of the tropopause accurately, the MTP does show that the GEOS-1 analyses clearly *overestimate* the tropopause temperature. Comparison of GEOS-1 data with ER-2 in situ temperatures (not shown) indicate that, though GEOS-1 tropopause positions are correct to about .5 km, temperatures are too warm by about 3K.

With the possible exception of the HAC2s observed during the 951220 and 960217 flights near the south end of the flight track, the HACs were clearly quite far from

any active convection reaching the tropopause (*cf* Figure 3). Thus, it is useful to use trajectory analysis to investigate the thermal and convective history of the air parcels in which the HACs are embedded. In particular, we would like to know: (1) what is the relationship of the HACs to convective outflow?; (2) what is the lifetime of HAC that originate or appear to originate from convective outflow?; (3) what is the relationship of the observed HACs to the temperature history?; (4) Can HACs form from clear, uncloudy air; and (5) if so, how? The convective history of the air establishes whether convective systems high enough to reach the levels at which the HACs were observed affected the air parcels, and if so, when that occurred. The thermal history of the air indicates whether a cloud formed by convective injection can survive until the observation point is reached. It can also provide an indicator, through the minimum temperature encountered during the trajectory history, of what the maximum water vapor content of the air might be, and how close the large scale temperatures at the point of observation are to saturation. If, for example, a parcel has encountered substantially colder large scale temperatures in the past than those observed currently, one would not expect to observe large or synoptical scale clouds. Clouds should either be absent, or require mesoscale cooling to form.

4. Trajectory Analysis Method

Based on Figure 2, all the observed HAC have a significant presence on the 380K surface; thus, we will seek to answer the science questions posed above by calculating 10-day adiabatic back trajectories on that isentropic surface from a dense set of points

along and adjacent to the DC-8 flight track for each of the four flights. It should be noted that the *actual* 380K surface in the tropics at this time of year is about .5 km lower than the GEOS-1 analyses indicate. This is due to the fact that the limited vertical resolution of the analyses (about 1.5 at these levels) “rounds out” the sharp tropical tropopause, resulting in tropopause temperatures (and potential temperatures) that are systematically too high. Tropical tropopause locations (in pressure), however, are very reasonable. Since the clouds are located near the tropopause, it is best to choose an isentropic surface that is also near the GEOS-1 model tropopause.

The trajectory model used is that of *Schoeberl et al.* [1995], and it is coupled to the GSFC GEOS-1 analysis fields, available at 6 hour intervals. The 10-day limit is chosen because radiative heating and cooling in the tropical tropopause region place limitations on the length of time over which a trajectory calculation is valid. *Jensen et al.*, [1999] estimate that the radiative heating rate at the tropical tropopause is about 0.5K/day. Such a heating rate changes the potential temperature of a parcel by about 10K in 10 days. Since this is the approximate potential temperature change over the depth of the observed HACs, as shown by Figure 2, this is a reasonable vertical “spread of uncertainty” in the trajectories.

To evaluate the convective history, the temperature along each trajectory is compared to 3-hourly brightness temperatures from the 10.5 micron channel of the GMS-5 and GOES-7 (December, 1995) and GOES-9 (February, 1996) geostationary meteorological satellites. If the brightness temperature is less than or equal to the parcel temperature along the trajectory, we assume that the convective cloud has reached

high enough (and attained temperatures cold enough – temperature decreases with altitude) to interact with the parcel on the trajectory. This interaction would take the form of a convective anvil outflow into that air parcel. This approach almost certainly underestimates the ability of convective clouds to "influence" the back trajectories, since it is implicitly assumed that the brightness temperature is equal to the cloud top temperature (The cloud top temperature then determines the cloud top altitude via the analysis temperature profile). In fact, brightness temperature will always be somewhat warmer than the actual cloud top temperature, since some radiation from warmer underlying surfaces will penetrate through the very top of the cloud. In the case of thinning anvils, these underlying surfaces might be the ocean surface. For somewhat thicker anvils, the fact that the cloud top is not a solid ice surface (!) means that the brightness temperature is a measure of the temperature some distance into the cloud (which will be warmer than the cloud top temperature) The difference between cloud top temperature and brightness temperature is negligible for the core of the convective cloud [Liou *et al.*, 1990], but may be significant for the outer parts of the anvil. We use a simple approach to correct for this effect. The size of typical convective systems is on the order of 1 degree by 1 degree. Since tropical anvils have a roughly constant cloud top altitude [Danielsen, 1993], we assume that the brightness temperature at any position in a typical 1 by 1 degree square is actually equal to the *minimum* brightness temperature in that square, provided that there is significant cloud at that position. We define significant cloud as being a brightness temperature of about 240K, equivalent to an optically thick cloud with an altitude of about 9 km (about the lowest level for

which significant detraining occurs in large tropical convective systems – *Houze et al.*, [1989]). This scheme has the effect of "flattening" the brightness temperatures for large convective systems, and the result is a more realistic picture of the actual cloud top temperature than the brightness temperatures themselves.

Despite this correction we may still obtain an underestimate of the impact of convection on large scale trajectories. Diabatic radiative heating rates are, in fact, not large at the tropical tropopause (0.5K per day, as indicated above). However, these calculations apply to clear sky situations only. *Ackerman et al.*, [1988] have shown that moderately thick anvils have heating rates on the order of 18-20K per day, effectively raising the *potential temperature* of the anvil at a rate of about 34K per day (or 1.4K per hour). With a lifetime of 5 hours [*Sherwood*, 1999], this implies a potential temperature change of about 7K. Such a change in the potential temperature of the anvil will have one of three effects on the anvil and its environment: (1) a change in the temperature of the anvil (and the environmental temperature profile) with no change in the anvil altitude; (2) a change in the anvil altitude with no change in the anvil temperature or environmental temperature profile; and (3) some combination of (1) and (2). A scale analysis by *Ackerman et al.*, [1988] suggests that most of this heating and potential temperature change in the anvil will exhibit itself as anvil ascent in an environment in which the temperature profile remains roughly unchanged. Thus, the anvil will actually cool as it follows the decreasing temperatures of the existing temperature profile ([*Sherwood*, 1999]). Typically, lapse rates in the tropical upper troposphere between 150hPa and the tropopause vary from about 7K/km to 4K/km. This is equivalent to

positive vertical potential temperature gradients between 5K/km and 10.5K/km. Thus, a potential temperature *increase* of 7K for an anvil would imply a temperature *decrease* between 10K (for the 7K/km temperature lapse rate) and 3K (for the 4K/km lapse rate). Thus, convective clouds that initially reach somewhat lower altitudes and warmer temperatures than the parcels along a back trajectory can, due to rapid anvil heating during the short period in which they are optically thick, actually reach those back trajectory parcels and inject air and ice particles into them. In view of this radiatively forced ascent, we shall assume that convective interaction occurs along a trajectory when the brightness temperature is less than 10K warmer than the parcel temperature. Based on observed lapse rates, this should represent the largest amount of convective influence that we would expect along a trajectory.

5. Results

Figure 3 (a-d) shows a sample of the temperature history of 10 day back trajectories on the 380K surface (a reasonable potential temperature level for the tropopause and for the observed HACs in Figure 2) superimposed on the brightness temperature images at the approximate flight times of the four flights. The gray scale along the trajectories shows the temperature history of each parcel. Three points are to be noted. First, the bulk of the trajectories come from the north and west. Some of these originate either in the Indonesian convective region or in the cold center near $[150^{\circ}E, 12^{\circ}N]$ (see Figure 1), and are effectively traveling around the monsoon anticyclone; others originate at the south edge of the subtropical jet. For all of the calculated trajectories, though, there

is no evidence of any that came from a latitude greater than 35°N . Second, for three of the flights, 951215, 951220 and 960217, a significant number of trajectories from the south end of the flight track come from the general direction of the SPCZ. Third, the temperature history for these two basic types of trajectories (from the north and west vs from the south), is distinctly different. For the northern trajectories, temperatures immediately prior to the time of observation were generally significantly warmer than at the time of observation, in some cases as much as 20K warmer. For the southern trajectories, temperature fluctuations during the 10 day history are much smaller (except for the strong penetrations into the southern summer midlatitudes in the flights of 951220 and 960217 depicted in Figures 3b and 3d). The implication is that the air in all the HACs on 951215 and 960213 and the northern HACs on 951220 has been too warm in the recent past to support clouds, implying fairly recent *in situ* formation.

Figure 4(a,b) shows the most important convective history parameter, namely the time since the most recent convection (if less than 10 days) experienced by a 1 by 1 degree grid of parcels at 12 GMT on 951220. This was calculated by using the convective history method described in section 4 using 10 day back trajectories. For Figure 4a, post-convective anvil heating (section 4) was assumed, while for Figure 4b, no such anvil heating was assumed. As can be seen by the large areas in white, most of the parcels have not been influenced by convection in the most recent 10 day period, reflecting the fact that the amount of area in the tropics occupied by thick anvils reaching the tropopause is fairly small. In fact, Figure 4a probably represents an upper limit, given the assumptions about the lapse rate described in section 4. For

Figure 4(a

those parcels north of the equator that have been influenced by convection, the time since that occurred is consistently quite long – a week or more. This reflects the fact that this convection occurred in the Indonesian region, and a significant amount of time has elapsed as the parcels have traveled around the anticyclone. Times are shorter for parcels from the SPCZ region. Notably, however, these times are often still several days in length, although the shortest times (within the SPCZ itself) are well under a day. Though the assumptions for convective interaction between Figures 4a and 4b are quite different, the basic picture of convective influence in the west central tropical Pacific is still qualitatively the same, with scattered convective influence north of the equator, and a fairly solid mass in the convective SPCZ region. Both figures have streamers of convective influence extending northeastward toward the flight track from the SPCZ. The northernmost streamer represents air that has been influenced by convection about 3 - 4 days previous, while the southern streamer (present only in Figure 4a) has been influenced by convection within the previous 12 hours. The positional uncertainty in the trajectory calculations is probably too large to establish whether the cloud in Figure 2c between $-3^{\circ}S$ and $-7^{\circ}S$ is associated with the “older” or “younger” streamer. The flight of 951220 had the largest differences in convective history between the anvil heating and no anvil heating calculations. For other flights where temperature and convective history indicated that the cirrus cloud was the remains of a convective anvil, anvil ages calculated by the two methods were similar.

Figure 5(a-d) shows a summary of the temperature and convective history of the parcels along the flight track for each flight, plotted as a function of latitude (since

Figure 5(a

all the flights were largely north-south). The demarcation between points whose back trajectories lead to the north and west and those whose back trajectories lead to the south is shown as a black slash on the axis, while the locations of observed HACs at 380K is denoted by heavy horizontal black lines. Though HACs were observed on all flights, each flight does have a different pattern of the types of HACs and where they were observed. Thus, we will discuss the temperature and convective history of the air parcels sampled by each flight in turn. Figure 5 does not show the time since the most recent convection, only whether convective influence occurred. These “convective lifetimes” will be indicated in the text where appropriate.

Perhaps the simplest flight to understand is that of 960217, where HAC2s at 380K were confined to the region south of about $2.5^{\circ}S$ (Figure 2 g and h). This location corresponds roughly to the demarcation between southerly and northerly trajectories (Figure 3d) and the region of convective influence in the past 10 days. There are two regions of convective influence shown in Figure 5d, both south of $3^{\circ}S$. The northern one ($3 - 5^{\circ}S$) has a variety of convective sources with a minimum age of about 3 days, and the southern one ($7 - 11^{\circ}S$) has a single source with an age of about 1.6 days. In both cases, the source of convective influence is the SPCZ. Notably, the gap in convective influence at $6^{\circ}S$ in Figure 5d coincides roughly with a minimum in the total scattering ratio at $6 - 7^{\circ}S$ between two maxima (Figures 2g and 2h). Figure 5d shows that, in the southern area of convective influence, a gradual cooling followed the convective injection, with minimum temperatures only slightly below those at the time of observation reached .1 to .3 days before the time of observation. Given the uncertainties in the trajectories

and the temperatures, one would conclude that the residual cloud has probably not evaporated during that last .1 to .3 days of slight warming. Just north of the southern convective influence region, Figure 5d indicates a sudden jump to very high maximum values in the temperature history. This is due to the midlatitude origin of these air masses and their passage through the SPCZ region during periods when convective activity did not reach high enough altitudes. The absence of convective influence, however, does not imply that no cloud should be observed. In fact, throughout the $5 - 8^{\circ}\text{S}$ region, the temperature history indicates continual cooling with a minimum attained less than 0.5 days before the time of observation. Again, the presence of cloud is not inconsistent with this temperature history – the warming at the end of this history is too small and occurs over too short a time to argue conclusively that the cloud that may have formed during the prolonged cooling has evaporated. The temperature history in the region between 3°S and 5°S is more variable, with convective influence from a variety of sources. The parcel most recently influenced by convection, however, is at about 4.5°S (3 days). Following that injection it then cooled to about 192K about 1.2 days prior to the time of observation, followed by warming to 196K and subsequent cooling to 194K. In fact, all the parcels in this $3 - 5^{\circ}\text{S}$ region underwent significant warming (2-3K) over a greater than 1 day period prior to the time of observation. This suggests that a 2-3K warming may be insufficient to evaporate all the ice crystals injected by the convection and condensed by the cooling following that convection.

The back trajectories for 960217 from locations north of $2 - 3^{\circ}\text{S}$ are all essentially from the north and west, passing at least partway around the monsoon anticyclone

(Figure 3d). North of $3^{\circ}N$, the parcels have experienced temperatures at least 2K colder than the current temperature more than a week prior to the time of observation. In fact, north of $3^{\circ}N$ the coldest temperatures in the 380K 10 day parcel histories are generally the coldest of all the four flights. North of $6^{\circ}N$, the 380K temperatures at the time of observation on this flight are generally the warmest of all the four flights. Arguably, then, the cold temperatures in the parcel histories have "wrung out" significant amounts of water vapor through condensation and subsequent sedimentation. The cooling of the parcels as they move southward towards the equator in the eastern branch of the anticyclone is insufficient to form clouds. There is a region between $3^{\circ}S$ and $3^{\circ}N$ where the coldest temperatures are very close to the temperatures at the observation point. This is not inconsistent with the cloud observations, though, as there is some observed cloud at $2^{\circ}N$ (Figure 2g).

The flight of 960213 (Figure 2 e and f; Figure 3c; Figure 5c) is a complete contrast to the 960217 flight in that it has only HAC1 type clouds. As Figure 3c shows, all the back trajectories are northerly, having traveled partway around the northern hemisphere anticyclone depicted in Figure 1b. This flight actually has the largest area of convective influence (Figure 4c) within the past 10 days, probably because a larger proportion of the trajectories are from the western Pacific convective region (Figure 3c). However, as Figure 5c shows, all this convectively influenced air cools somewhat after the convection, as indicated by the asterisks in the figure. This would be expected to lead to further ice crystal growth, subsequent sedimentation and consequent loss of water vapor. The air then warms by between 5K and 15K, presumably due to descent as it travels northward

into the subtropical jet region prior to arriving in the area of observation. Any residual ice crystals from the convective outflow would thus almost certainly evaporate. Effectively, all the “memory” of the convective injection temperatures has been “lost” as the parcels cool and warm following the convection. In fact, there is little apparent spatial relationship between the HACs and the region of convective influence. Overall, except for a region between $3 - 4^{\circ}N$, the minimum parcel temperatures in the history are all substantially lower than the current temperatures. Arguably, then, one would not expect formation of clouds on a large scale and, indeed, we did not observe the extended continuous sheets seen by *Winker and Trepte* [1998]. Rather, one sees sloping structures that actually occupy a mesoscale region at any given altitude or potential temperature surface. Section 6 below discusses how gravity and inertia-gravity waves might generate these structures.

The two February, 1996 flights each contain one of the two HAC types, whereas the December, 1995 flights have both cloud types. Turning first to 951220 (Figure 2 c-d), we note a HAC2 cloud south of about $2.5^{\circ}S$, and HAC1 between $8^{\circ}N$ and $16^{\circ}N$. On the 380K surface, the air parcels along the flight track have the same basic pattern of trajectory origins (Figure 3b) as for the 960217 flight, with air south of the equator having basically southerly trajectories, and air north of the equator, northerly. Convective influence within the past 10 days is strongest (in terms of proportion of parcels influenced) south of about $2.5^{\circ}S$ (Figure 4b), with two regions about 0.5 days ($6^{\circ}S$ to $4.5^{\circ}S$) and 3.5 days ($3^{\circ}S$) old. The temperature history for the southernmost of these two regions (Figure 5b) indicates roughly constant temperature between

the convective injection and the time of observation – certainly consistent with the continued presence of cloud. For the region near $3^{\circ}S$, the air cools perhaps 2K, and then warms slightly for about 0.5 days prior to observation. Though not as clear cut as for the southernmost region of convective influence, these temperature variations are not large enough (when uncertainty is taken into account) to rule out the possibility of convective outflow cloud. Notably there is no convective influence north of $2.5^{circ}S$, which is the northern edge of the observed cloud in Figure 2c. This further reinforces the interpretation that the HAC2 cloud south of $2.5^{\circ}S$ is a residual from convective outflow. Convective influence is also found north of about $7^{\circ}N$. As for 960213, however, this convective injection has occurred between 7 and 10 days prior to the time of observation, and is followed by the same substantial warming as the air parcels flow northward into the subtropical jet. Thus, as for 960213, the observed clouds between $8 - 16^{\circ}N$ are clearly not due to convective outflow. Unlike 960213, however, the coldest temperatures in the parcel history for much of the HAC1 region ($10.5 - 14^{\circ}N$ – Figure 5b) are quite close to the current temperatures, and these temperatures occur within 0.5 days of the time of observation. Thus, in this case it is quite plausible that the large scale cooling associated with the equatorward air motion could be responsible for much of the HAC1 between $8 - 16^{\circ}N$. In fact, the larger extent of the cloud at the 380K and 390K potential temperature surfaces (about 600 km) vis a vis the much narrower sloping clouds on 960213 is consistent with this scenario. The temperature history along other parts of the flight track that have no cloud is reasonably consistent with the cloud observations. Between $2.5^{\circ}S$ (the northern edge of the observed cloud) and about $4^{\circ}N$

the parcels experienced significantly colder than current temperatures between 1 and 7 days prior to the time of observation. The implication is that water vapor has been removed from the air parcels by condensation due to cold temperatures. Subsequent warming of 3-4K would then evaporate any remaining ice crystals.

The flight of 951215 (Figure 5a) is the most ambiguous of the four in that this is the only case where HAC2s are found in the northern portion of the flight track. As in the previous flights, parcels at the southern end (south of about $3^{\circ}S$ in this case) have sufficient southerly trajectories to reach the mean position of the SPCZ (Figure 3a). There are regions south of $3^{\circ}S$ where convection has influenced the air parcels within the past 5-7 days. Unlike 960217 and 951220, however, where HAC2s were observed in these convectively influenced regions, there is no evidence of cloud south of $3^{\circ}N$. This absence of cloud is consistent with the temperature history, which shows cooling following the convective injection to temperatures 3-4K below the temperature at the time of observation. These minimum temperatures occurred 2 days prior to the time of observation. These cold temperatures may well have removed water vapor above the saturation mixing ratio. The air subsequently warmed by 3-4K until the time of observation; this warming presumably evaporated any left over ice crystals. North of $3^{\circ}S$, convective influence is very minimal. Moreover, the temperature histories indicate very little potential for large scale cloud formation along the flight track between $3^{\circ}S$ and $8^{\circ}N$. This is because there is significant warming of 3-4K in the 2-5 days prior to the time of observation. As suggested above, the earlier cold temperatures would arguably remove sufficient water vapor to make the air too dry to condense at temperatures at

the observation time. However, there is, in fact, a quite thick HAC2 between $4 - 7.5^\circ N$, entirely inconsistent with these expectations. There is also significant cloud north of $8^\circ N$, both of the thick HAC2 variety and the thinner, sloping HAC1 (Figure 2 a and b). In contrast to the region between $4 - 7.5^\circ N$, temperatures at the observation time between $8 - 15^\circ N$ are actually quite close to the coldest in the entire parcel history, indicating that this region does have the potential for producing cloud. The flight of 951215 differs from the other 3 in one important respect, which may account for the thick clouds observed between $4 - 12^\circ N$ - the northern ITCZ at around $8 - 10^\circ N$, west of the flight track was unusually active. This has no direct impact in terms of convective blowoff, since tropopause winds are westward, away from the flight region. However, these convective systems may generate high amplitude gravity waves that produce significant mesoscale temperature variations, temperature variations that might account for the observed HAC2s.

In summary, the four case studies show that the presence or absence of HACs, and their morphology, is at least qualitatively consistent with the convective and temperature history of the air parcels. For the portions of the flight tracks where the air parcels had northerly back trajectories, convective injection, if any, occurred on the order of a week prior to the observation time. This convective injection was usually followed by: (1) further cooling of a few degrees; (2) significant warming to temperatures 5-10K or more above those at the time of observation; and (3) cooling to the current temperature. The major implication is that any convective blowoff in this northerly back trajectory region has almost certainly evaporated prior to the observation time.

Thus, clouds in this region are condensed from clear air, and, except for 951215, are thin, sloping structures with a low total scattering ratio (HAC1). The extent of these clouds is, for the most part consistent with the temperature history. In other words, it is qualitatively consistent with the notion that cold temperatures earlier in a parcel's history can condense and remove sufficient moisture so that temperatures must get at least as cold again to form cloud. When the temperatures at the time of observation are close to the coldest in the air parcel history (951220 and 951215) cloud coverage is fairly extensive. Otherwise, there is either no (960217) or very limited (960213) coverage. Where cloud is observed in contradiction to the temperature history, there is either limited indirect evidence (951215) or very direct evidence (960213 – see section 6 below) for mesoscale motions on the scale of the observed HACs.

Those portions of the flight tracks where the air parcels had a southerly back trajectory always had significant, fairly recent convective influence from the SPCZ. Three of the flights (951215, 951220, and 960217) had such convective influence, and in two cases fairly thick HAC2s were observed. It is likely that these clouds were residual ice crystals from convective blowoff – that is, anvils that had been able to maintain themselves because of either slow cooling or only minimal warming. No cloud was observed in the southerly trajectory portion of the 951215 flight, due to significant warming between the time of convective injection and the observation time. If these clouds are indeed convective blowoff, it implies a significant lifetime for some of these anvils, ranging from 0.5 days (951220) to 3 days (960217) in the cases considered here. Such long lifetimes are implied by the recent work of *Sherwood*, [1999], who has shown

that the formation and radiative rise of convective anvils essentially advects the mean moisture gradient upward. This advection supplies moisture to the tropopause region and, in the absence of significant synoptically imposed downward motion, can allow a thin cirrus cloud to maintain itself. It should be noted that, due to sedimentation, the ice particles in the HAC are unlikely to be the same ones that were presented right after convection occurred. Thus, a “3 day lifetime” for an outflow cirrus cloud really means that the relevant air parcel has been occupied by cloud continuously since the formation of an anvil by convection.

Two shortcomings of the GEOS-1 analyses, and their impact on these conclusions, require some discussion: (1) the GEOS-1 analyses overestimate the tropopause temperature by about 3K; and (2) these analyses fail to include significant mesoscale temperature fluctuations. This is evident from MTP measurements (Fig 2) and from explicit comparisons with in situ aircraft data. The impact of these limitations of the analyses on our conclusions are probably small. First, the overestimate of tropopause temperatures appears to be due to the “rounding” of the sharp tropical tropopause by the limited (1 km) vertical resolution of the analyses. Since most of the trajectories are confined to the tropical regions, with similar tropopause “sharpness,” the temperature error is probably systematic. Since the conclusions of the trajectory analyses are based on comparative rather than absolute temperatures, a largely systematic error should not impact them. Mesoscale temperature fluctuations are clearly important, and, for sufficiently long time scales, *Jensen et al.*, [1996] have shown a significant impact on water vapor content. However, *Jensen et al.*, [1996] also show that the shortest scales

may have very little impact because the very rapid cooling rates produce large numbers of small particles. These are not able to sediment prior to the air parcel returning to equilibrium, reheating, and reevaporating the small particles.

6. Generation of high altitude cirrus by Gravity Waves – a Case Study

The flight of 960213 is unique in two ways. First, both the narrow mesoscale extent of the clouds and the temperature history indicate that mesoscale temperature or water vapor fluctuations are responsible for the clouds. Second, an ER-2 flight of similar length to the DC-8 flight occurred along the same horizontal path. This flight was centered in time about 16 hours after the midpoint of the DC-8 flight. Hence, *in situ* sampling was available where the large scale conditions (if not the mesoscale conditions) were very similar to those during the DC-8 flight. Presumably the character of the mesoscale motions during the two flights, if not the exact placement, would be similar.

The ER-2 flew directly south to about $1^{\circ}S$ on the $155^{\circ}W$ meridian at about 18.5-19.5 km. Upon reaching $1^{\circ}S$, the aircraft turned around and descended to 15.5 km, ascending immediately thereafter, thus providing two vertical profiles spanning about 4 degrees of latitude around the equator. With the exception of the profiles at the equator, the altitude of the ER-2 measurements is significantly higher than the region of HACs near the tropopause. However, since the vertical stratification at ER-2 altitudes is comparable to that in the tropopause region, we expect the nature of the

mesoscale motions at the tropopause (e.g., gravity waves) to be more nearly similar to the mesoscale motions at ER-2 altitudes than those at DC-8 altitudes, where the atmosphere is close to adiabatic. The available relevant ER-2 measurements were: *in situ* temperature, pressure and winds (Meteorological Measurement System, or MMS – *Gaines et al.*, [1992]), vertical temperature profiles in an 8 km deep curtain centered at the aircraft altitude (Microwave Temperature Profiler, or MTP – *Gary*, [1989]), and water vapor from a Lyman- α hygrometer (*Hintsa et al.*, [1998]). The critical questions are: (1) is there evidence of mesoscale temperature fluctuations with a structure resembling that of the observed HAC1s?; (2) what is the mechanism responsible for these fluctuations?; and (3) is there evidence of water vapor at or near saturation in the *in situ* measurements near the tropopause?

Figure 6 shows an ER-2 MTP latitude-height temperature curtain, superimposed on the southbound DC-8 total scattering ratios taken about 16 hours before the ER-2 curtain. Note that the MTP provides its highest vertical resolution near the aircraft (in this case the ER-2), which means that phenomena of short vertical scales will "die off" away from the aircraft due solely to instrumental considerations. Also, as noted above in the discussion of Figure 2, overall accuracy decreases as distance from the aircraft increases. Thus, the increase in tropopause region temperatures towards the equator exhibited in Figure 6 is almost certainly an instrumental effect due to the gradual increase in the ER-2's altitude. Between about $10 - 18^{\circ}N$ at about 19km there are upward and northward sloping structures. This sloping feature in the temperature perturbations is a hallmark of atmospheric internal gravity waves [*Holton*, 1972, pp.

Figure 6

172-176], though this this feature does not by itself prove that the phenomena are gravity waves. To do that, one must examine the other meteorological variables (*e.g.* horizontal winds), analyze their amplitude and phase relationships, and compare those relationships with theory. This is done in Figure 7, which shows the *in situ* ER-2 meteorological variables, bandpass filtered to show only variance with the characteristic horizontal wavelength of these structures (apparent wavelength between 180 and 500km). Of particular interest are the three temperature "waves" with temperature peaks at $12^\circ N$, $14.8^\circ N$, and $17^\circ N$ in Figure 7 (top panel). These waves coincide with the sloping structures described in Figure 6, as can be seen by comparison with Figure 7 (bottom panel) which shows the ER-2 MTP temperature curtain in contour form. The gravity wave properties based on the observed meteorological variables in the $11 - 18^\circ N$ region depicted in this figure, and summarized in the first line of Table 1, clearly show that the temperature structures are northward and upward propagating gravity waves. This is demonstrated in three ways. First, the quadrature (90 degrees out of phase) relationship between T and V is consistent with that expected for gravity waves. Secondly, hydrostatic gravity waves in a slowly varying environment exhibit a "universal" relationship between the temperature and horizontal wind perturbations described as:

$$\delta|V| = N\delta|Z| \quad (1)$$

where $\delta|Z|$ is the vertical displacement, uniquely related to the temperature perturbation and N is the buoyancy frequency. As Table 1 shows, the observed meridional wind

Figure 7

perturbation is very close to that derived from the temperature perturbation. Third, Figure 7 (top and middle panels) show that V and dT/dz (the vertical temperature gradient) are negatively correlated; this, along with the slope of the structures indicates that the phase propagation is downward and northward with respect to the flow, an indicator of upward and northward group velocity. It should be noted that though the zonal wind perturbation is not insignificant, it is smaller than the meridional wind perturbation and lacks any coherence with T and dT/dz ; this shows that the wave has little if any zonal component to either the horizontal wavenumber vector or the horizontal propagation direction. It is clear from Figure 6, however, that these gravity waves are not a candidate for explaining the observed HAC1s – the upward and northward slope of the gravity waves is simply too steep. Also, that slope is unlikely to have been significantly less steep at the lower altitudes from which the waves have propagated, since the meridional wind and stability that govern the vertical wavelength are substantially the same near the tropopause.

However, Figures 6 and 7 do show evidence for longer wavelengths than the 250-300km structures described above. Figure 8 shows meteorological variables that are low-pass filtered to exclude the wavelengths depicted in Figure 7 (and all those that are shorter as well). In the case of V , the mean value (for the flight leg between $0 - 20^\circ N$) and the trend have been removed as well. At these scales, we also see an upward and northward slope, except that the slope is much shallower than for the shorter waves. As before, there is a clear quadrature (90 degree phase) relationship between V and T that suggests that the sloping structure is a gravity wave, and an

Figure 8

amplitude relationship between V and $N\delta|Z|$ that confirms it (second line of Table 1). Though there is significant zonal wind variance, its horizontal wavelength is generally different than that of the meridional wind. To the extent that there is coherence, it is roughly in quadrature with the meridional wind, possibly suggesting an inertia-gravity wave. However, the inferred (from the vertical wavelength) Doppler shifted frequency is 3 times the Coriolis parameter, which means that this wave is most likely a pure, hydrostatic, planar gravity wave. This wave, or one like it, is a more likely candidate for explaining the sloping HAC1s between $3 - 14^\circ N$, though the wave slope (or ratio of horizontal to vertical wavenumbers) is about 1.5 times that of the slope of the HAC1s (cf third line of Table 1).

Nevertheless, gravity waves with slopes and horizontal wavelengths roughly comparable to those of the clouds are clearly present within 10-20 hours of the cloud observations. Moreover, those gravity waves have propagated upward, indicating that they have been at the tropopause at one time even though they are observed at 19 km. Their temperature amplitudes are substantial, 1.3K for the short waves and 2.4K for the long waves, implying saturation mixing ratio changes of .9 ppmv and 1.8 ppmv respectively from the background. The implication is that mesoscale gravity waves alone are capable of producing subvisible cirrus clouds at the tropopause. Thus, as *Jensen et al.*, [1996] and *Potter et al.*, [1995] indicate, these motions could well play a role in dehydrating the air in the tropopause region prior to its entry into the stratosphere. Another key parameter is the intrinsic period of the gravity waves (the period as seen by an air parcel flowing through the waves). This is important because it governs the

cooling rate experienced by the air, and hence the size of the ice crystals that could be formed [Jensen *et al.*, 1996]. The periods of the three different waves illustrated in Figures 6, 7, and 8 (including the "wave" determined by the cloud structure, third line of Table 1) vary from 4 to about 40 hours. Based on Jensen *et al.*'s [1996] work, the two longer period waves would certainly imply cooling rates that are sufficiently slow to produce large enough ice crystals for fallout and dehydration in the time scale of the wave.

Further evidence that gravity waves could produce HACs is shown in Figure 9 (a and b). These are vertical profiles, with the descent (Figure 9a) extending from $1^{\circ}S$ to $0.5^{\circ}N$ while the ascent (Figure 9b) extends from $0.5 - 3^{\circ}N$. The most important feature of these profiles is that both are saturated near the tropopause, with the descent saturated from 16 to 16.7 km and the slightly colder ascent from 15.75 to 16.9 km. Though there were no instruments on the ER-2 for detecting cloud particles, pilot reports indicate very thin cirrus below the aircraft at $1.5^{\circ}N$, clearly consistent with the observations of saturation. The highest altitude of saturation is actually roughly consistent with the top of the HAC1 in Figure 2f at the lowest latitudes. A small region above the tropopause at 16.65 km (about 370K) in Figure 9b is clearly saturated, and this tropopause is actually fairly close to the top of the "blob-like" structure in Figure 2e at $3 - 5^{\circ}N$. Thus, a comparison of Figures 2 (e, f) and 9 suggests that the sloping portions of the HAC1s between $3 - 7^{\circ}N$ are actually just above the tropopause, with the blob-like structure below it.

Figure 9 (a

The role of gravity waves is illustrated by comparing the vertical profiles of

temperature and zonal wind in Figure 9 (a and b). There is a clear quadrature relationship in the vertical between these two meteorological quantities in the 16.5 to 19.5 km region, with a vertical wavelength of about 1.3 km. Also, as the fourth line of Table 1 shows, the quantitative relationship between the wind perturbation and $N\delta|Z|$ holds reasonably well. There is little coherence with the meridional wind, so this wave appears to be propagating roughly east-west. If an upward group velocity is assumed, the observations imply an eastward propagating gravity wave with a Doppler-shifted phase speed of about 5 ms^{-1} . Notably, this is comparable to the zonal wind amplitude of the wave, so the wave is actually quite close to breaking. At the tropopause, the wave's temperature perturbation will manifest itself partly as a rise and fall of the tropopause, resulting in a peak-to-peak variation of about 5K in the tropopause temperature (based on the temperature amplitude exhibited by the wave in the stratosphere). This implies a 2.5 ppmv peak-to-peak variation in saturation mixing ratio. It should be emphasized that the horizontal component of the aircraft's flight track is along (rather than perpendicular to) the wave's phase surfaces (the wave is propagating eastward while the aircraft is moving northward). Thus, the observed temperature differences between the two profiles at the tropopause due to the wave will be relatively small. These differences will be more a manifestation of latitudinal amplitude variation than of phase variation. The gravity wave does more than raise and cool the tropopause. As can be seen from Figure 9b, regions above the tropopause are also cooled (with respect to Figure 9a), and it is clear that there is a small region above the tropopause that is saturated.

The questions of how and where the of the observed gravity waves are generated is

important, since the answers to those questions will indicate how representative these observations are of the tropics as a whole. For the waves observed on the southward leg of the ER-2 flight, as well as those presumed responsible for the HAC1s observed by the DC-8 (first three lines of Table 1), we can calculate vertical and horizontal group velocities. The important point about these group velocities is that a crude backward trace to the tropopause level indicates a "point of origin" no further south than the equator. It is thus doubtful that these particular waves have their source at the SPCZ at $10^{\circ}S$ to $15^{\circ}S$. It is more likely that they are produced by imbalances in the flow caused by the upper tropospheric trough "digging" south towards the equator. That imbalances in the flow can produce gravity waves has been demonstrated by the modeling work of *O'Sullivan and Dunkerton*, [1995]. In their work, baroclinic development at midlatitudes produced inertia-modified gravity waves. At these latitudes, however, the coriolis parameter is sufficiently small that waves with similar frequencies are not substantially modified by the earth's rotation. If this imbalance in the flow is indeed the mechanism for generating these waves, one cannot conclude that similar waves are necessarily found at other longitudes (though there is strong evidence for inertia-gravity throughout the tropics – *Karoly*, [1996]). In particular, the very cold region near $130 - 175^{\circ}E$ and $10^{\circ}N$ (Figure 1a and 1b) does not typically experience strong digging troughs since it is "protected" by a strong and steady subtropical jet to the north.

The waves observed during the ER-2's dive at the equator may well have a different origin, though the absence of horizontal wavelength information limits the conclusions that can be drawn. The east-west propagation at least suggests the possibility that

equatorially trapped modes [Holton, 1975, pp 53 - 70] may play a role. If we assume that they are indeed equatorially trapped, forced gravity waves with a horizontal wavelength of less than 1500km, the dispersion relationship is essentially that of a pure gravity wave, with the latitudinal trapping proportional only to the vertical wavelength and the Doppler-shifted phase speed. Applying the relevant equations, these assumptions imply an equatorial trapping distance of about 5 degrees of latitude. Because of the equatorial trapping and the implied confinement of wave energy, a very substantial horizontal propagation distance is possible for these waves, suggesting that they may have been generated by convection in the western Pacific. The cloud data actually give an indication that there is longitudinal phase variation in the temperature, since the northbound leg at (Figure 2f) has significantly more cloud near the equator than the southbound leg *and* that leg is about 2.5 degrees (275km) further east.

7. Summary

This study has attempted to explain a given set of observations of thin cirrus clouds near the tropical tropopause with a view toward establishing their consistency with temperature and convective history, and our limited understanding of their formation mechanisms. It represents a complementary approach to that of [Newell et al., 1996], who compared DIAL observations of tropical cirrus clouds near the tropopause with analyzed vertical motion and relative humidity fields. The basic structures and nature of the clouds found during the TOTE/VOTE experiment were similar to the findings of from the space-based LITE experiment [Winker and Trepte 1998]. Two basic types of

clouds were found – extended laminar structures with horizontal extents up to 1000 km (HAC1 – similar to those shown in Plate 1 in [*Winker and Trepte* 1998]) and blobbier, thicker structures resembling convective clouds (HAC2 – similar to those shown in Plate 3 in [*Winker and Trepte* 1998]). Altitudes of these clouds were comparable to those in [*Winker and Trepte* 1998], though the apparent convective outflows (HAC2) were higher. The latter is to be expected, since the tropopause height in the tropical central Pacific in northern hemisphere winter is generally higher than the tropopause height in September (the time of [*Winker and Trepte* 1998] LITE observations). Comparing the TOTE/VOTE results with cirrus clouds observed during PEM-WEST by [Newell et al., 1996], it appears that Newell’s clouds most strongly resembled the HAC2’s in this study.

For the most part, the positions of the respective cloud types were consistent with their convective and temperature histories. The HAC2s (except for the 951215 case) were found in the southern portions of the Hawaii-equator flight track and were consistent with convective outflow from the Southern Pacific Convergent Zone near $10 - 15^{\circ}\text{S}$. In other words, back trajectories indicated encounters with significant convection in the previous 10 days *and* the temperature histories indicated that the air had either cooled since the convective encounter, or not warmed substantially. In one case where air had interacted with the SPCZ but no cloud was observed, the temperature history indicated substantial cooling after the convective encounter, followed by warming. Thus, the absence of cloud was entirely consistent with the temperature and convective history. For the observed convective outflow clouds, the time since the most recent convection ranged from 0.5 days to 3 days. This is an important finding, since it suggests that

convective outflow cirrus are not necessarily dissipated rapidly, but can remain in the tropopause region for extended periods of time. This increases their potential for impacting the radiation budget [Jensen *et al.*, 1996]. The implication of this finding in the context of the recent work of Sherwood, [1999] is that convective outflow can continue to rise and advect moisture into the tropopause region long after the convection has occurred. It also has potential importance to the dehydration issue. If the high supersaturation requirements (160%) for *in situ* ice nucleation of sulfate aerosols turns out to be a major barrier for the formation of *in situ* thin cirrus clouds at the tropopause [Heymsfield *et al.*, 1998], then the ability of convective outflow cirrus (where ice crystals are already present) to act as "seeds" for the formation of clouds may be very significant.

In the northern portion of the flight track, the thin laminar cirrus structures tended to dominate. Again, their presence or absence, as well as their structure, was for the most part consistent with temperature and convective histories. Air north of the equator (approximately) had its origin either in the northern hemisphere subtropical jet or in the convective regions of the western equatorial Pacific and the equatorial Indian ocean. In any case, all the air passed through or near the subtropical jet. Thus, it experienced temperatures upwards of 10 degrees warmer prior to being advected southward around the monsoon anticyclone (Figures 1a and b). The implication is that all the clouds were formed *in situ*, as any convective outflow would have evaporated prior to the time of observation. The structure and presence of the clouds were entirely consistent with temperature histories, with large scale thin cirrus observed only on the flight where temperatures at the observation time were close to the coldest observed

in the entire temperature history. The one time where clouds were observed in spite of the relatively warm synoptic scale temperatures, the clouds were strongly sloping, and were actually quite limited in scale at any given altitude. The cloud structures resembled upward and northward propagating gravity waves, and such gravity waves were in fact observed about 16 hours later at somewhat higher altitude with *in situ* ER-2 observations. The properties of these gravity waves indicated that they were potentially significant for dehydration, both from the point of view of the measured temperature perturbations and cooling rates. The temperature variations are several degrees peak-to-peak, implying a more than 2 ppmv variation in the saturation mixing ratio; this, in turn is about 70% of the typical value of the water vapor mixing ratio at the tropopause in northern hemisphere winter. The wave periods, varying from about 4 to 40 hours for cases isolated by the analysis, are long enough to allow large (10 microns or more) ice particles to form and fall out, thus effectively dehydrating the tropopause region [Jensen *et al.*, 1996]. If these waves are widespread in the tropics (and observations, e.g., Karoly *et al.*, [1996], indicate that they are) the implication is that they could effectively "ratchet down" the water vapor content in the tropopause region to values below those implied by synoptic scale analyses at the tropopause, or, for that matter, averages of tropopause cold point saturation mixing ratios from radiosondes [Dessler, 1998]. The success of Dessler's radiosonde average minimum saturation mixing ratio approach in producing a reasonable entry value for water vapor at the tropical tropopause suggests that this gravity wave mechanism may not work as well as [Jensen *et al.*'s, 1996] model suggests. One possibility is that particles require a certain level of

supersaturation to form [*Heymsfield et al.*, 1998]. Another possibility is that we have not considered that each air parcel experiences not just the single, long-period gravity waves we have isolated in Table 1, but the complete spectrum. Thus, even though the temperature perturbation in the ER-2 flight leg is dominated by the 1100 km wave (see Table 1), the cooling rate associated with the 4 hour wave is actually larger by a factor of 3 (one sixth the period, half the amplitude). A complete assessment of how or even whether [*Jensen et al's*, 1996] mechanism actually dehydrates air in the upper troposphere and lower stratosphere requires consideration of the entire spectrum of gravity waves, including some appropriate spectral model [*Bacmeister et al.*, 1996].

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Figure 1. Average 10.5 micron brightness temperature as observed by the GMS-5, GOES-7 and GOES-9 meteorological satellites during two periods of winter 1995-1996: (a) December 5-25, 1995; and (b) February 1-19, 1996. Average winds and temperatures

on the 85 hPa surface for both periods from GSFC GEOS-1 analyses are superimposed.

Figure 2. Total Scattering Ratio (TSR) at .6 microns for four flights during TOTE/VOTE as a function of latitude. Color fill shows temperatures observed by the Microwave Temperature Profiler (Gary, 1989). White lines are contours of $\log(\text{TSR})$, with contour intervals of .05. Heavy blue lines are isentropes from the GSFC GEOS-1 meteorological analyses. Dates of the flights are indicated on the title of each figure.

Figures a, c, e, and g are southbound legs. Figures b, d, f, and h are northbound.

Figure 3. Back trajectories on the 380K surface from selected points along the DC-8 flight track for the flights of 951215(a), 951220(b), 960213(c), and 960217(d). Temperature histories are color coded along the trajectories. The back trajectories are superimposed on the 10.5 micron brightness temperature images valid at the approximate flight times of the four flights.

Figure 4. Gray scale coded times since the most recent encounter with convection of 10 day back trajectories originating in the west central Pacific region at flight time on 951220: (a) assuming post-convective anvil heating as described in the text; (b) assuming no post-convective anvil heating.

Figure 5. Summary thermal histories of back trajectories on the 380K surface from points along the DC-8 flight track for the flights of 951215(a), 951220(b), 960213(c), and 960217(d). The solid line denotes the temperatures on the 380K surface at the time of observation by the DC-8 along the flight track. The diamonds denote the warmest temperatures along the trajectory. The asterisks and triangles denote the coldest temperatures reached before and after the time of this "warm point," respectively. Parcels that have been convectively influenced in the past 10 days are denoted by solid squares. For these parcels, only times AFTER the time of convective influence are considered in evaluating the "warm point" and the two surrounding "cold points."

Figure 6. Contours of $\log(\text{TSR})$ at intervals of .05 for the southbound leg of the DC-8 flight of 960213 (white contours) superimposed on MTP temperatures for the southbound leg of the ER-2 flight of 960213 about 16 hours later.

Figure 7. Meteorological variables for the southbound leg of the 960213 ER-2 flight: (a) *in situ* temperature (solid) and meridional wind (dotted) from the ER-2 MMS; (b) *in situ* vertical temperature gradient from the ER-2 MTP (solid) and zonal wind (dotted); (c) temperature contours from the ER-2 MTP. (a) and (b) are band passed to allow only wavelengths between 180 and 500 km. (c) is low-pass filtered to allow wavelengths longer than 180km.

Figure 8. As in Figure 8, except that (a) and (b) are now low-pass filtered to allow only wavelengths greater than 500 km. (c) is now the temperature DEVIATION from the mean vertical profile low pass filtered to include only wavelengths greater than 500km.

Figure 9. *in situ* water vapor, ice saturation mixing ratio, zonal wind, and temperature from the ER-2 descent and ascent near the equator on the flight of 960213. (a) Data from descent centered at about .4S; (b) data from subsequent ascent centered at about 1.5N.

Table 1. Times and locations of observed high altitude cloud types

High Altitude Cloud Type	Flight Date	Locations in Latitude
HAC 1	951215	$10^{\circ}N - 15^{\circ}N$
	951220	$6^{\circ}N - 16^{\circ}N$
	960213	$0^{\circ}N - 13^{\circ}N$
	960217	<i>none</i>
HAC 2	951215	$2^{\circ}N - 11^{\circ}N$
	951220	$7^{\circ}S - 3^{\circ}S$
	960213	<i>none</i>
	960217	$11^{\circ}S - 8^{\circ}N$

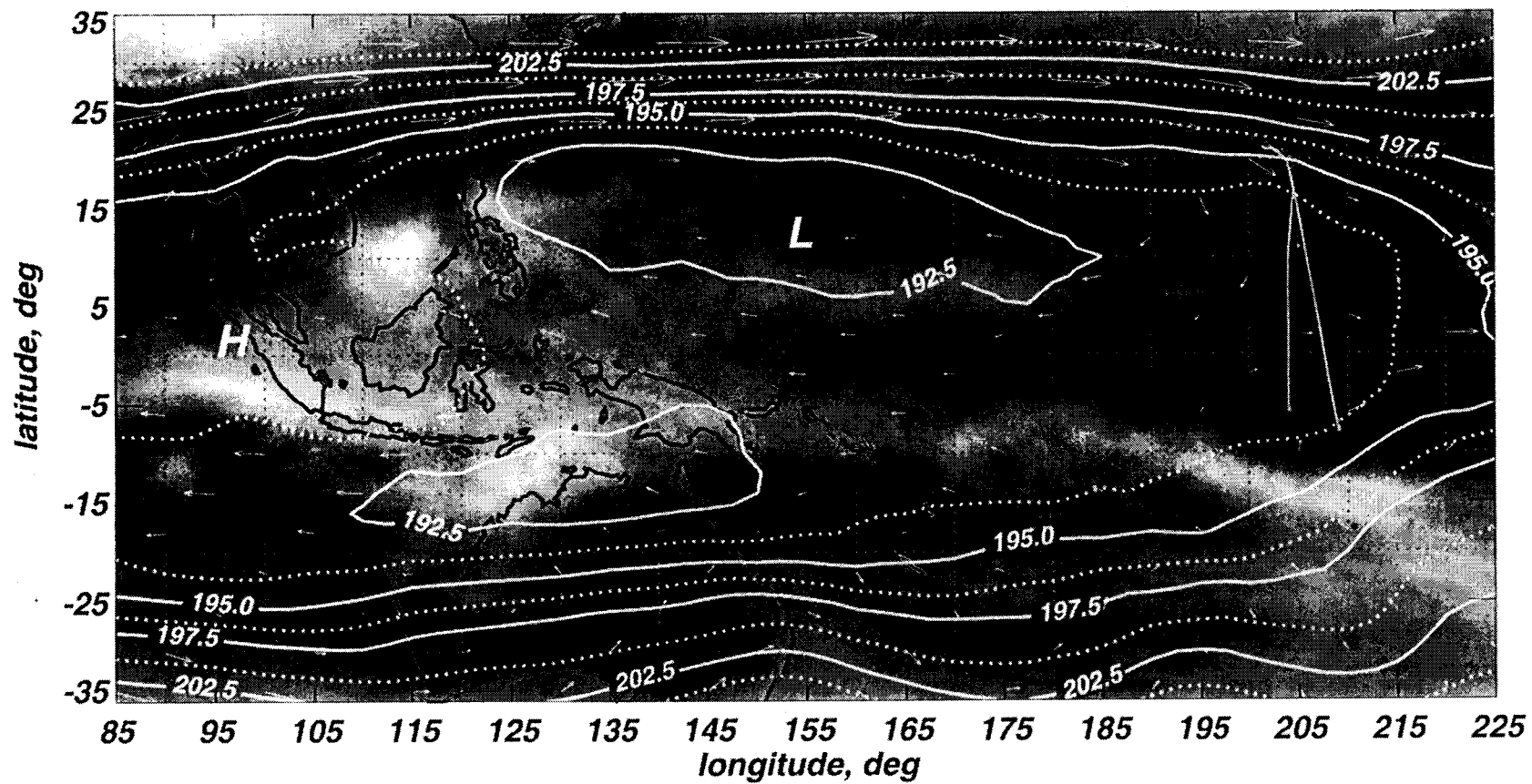


Fig. 1a

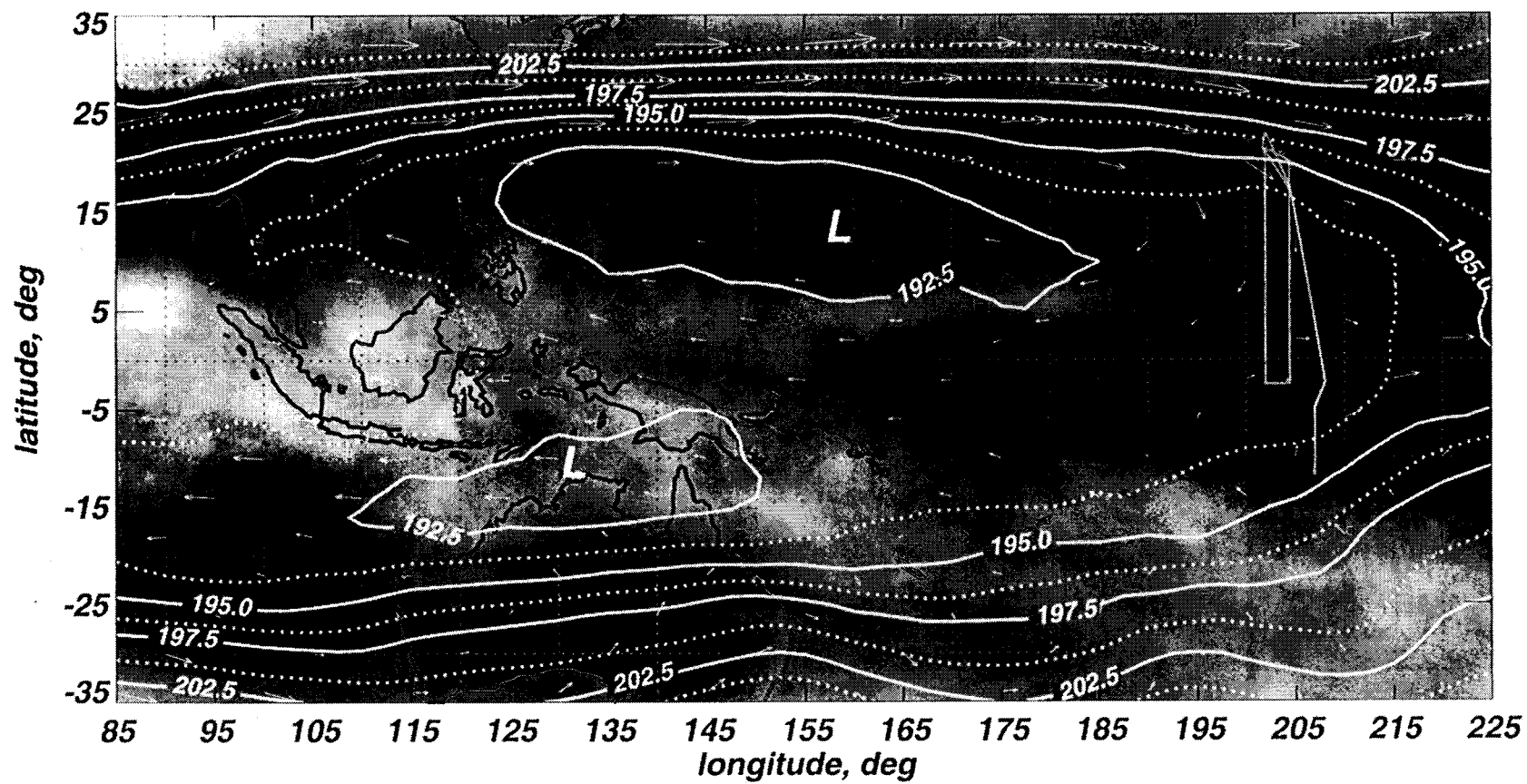


Fig. 1b

Fig 2a

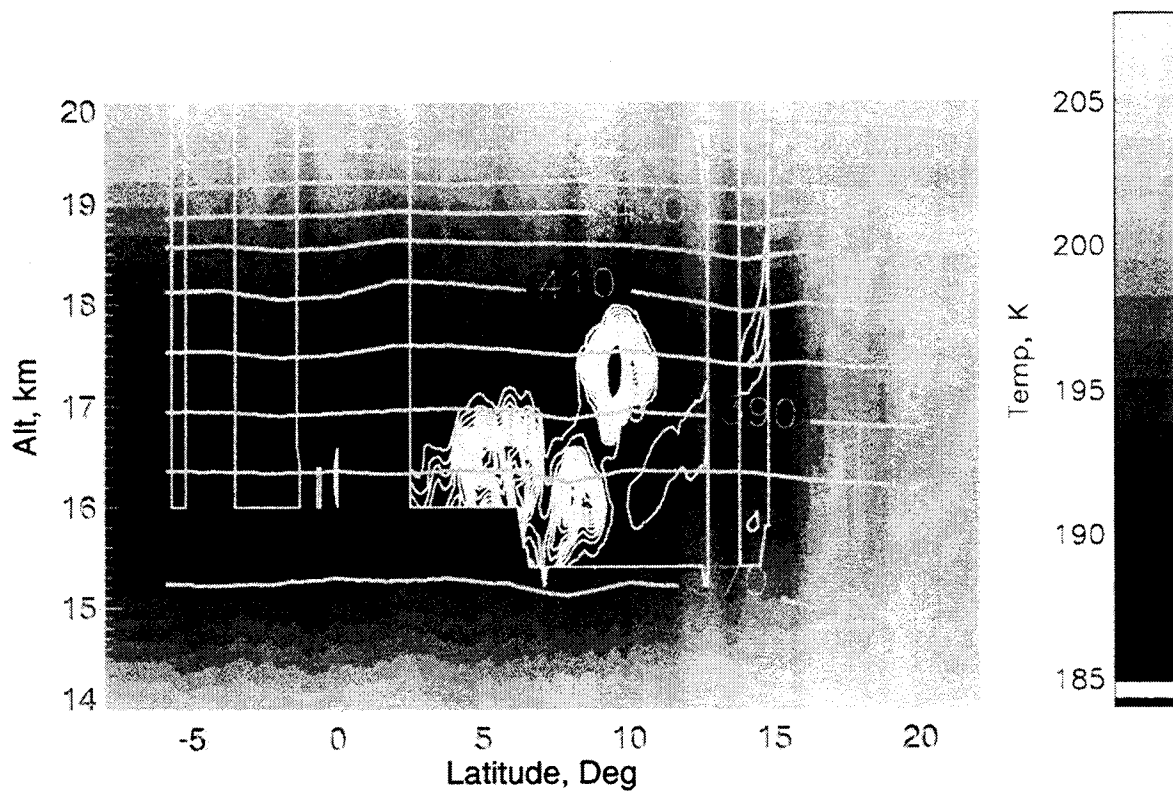


Fig 2b

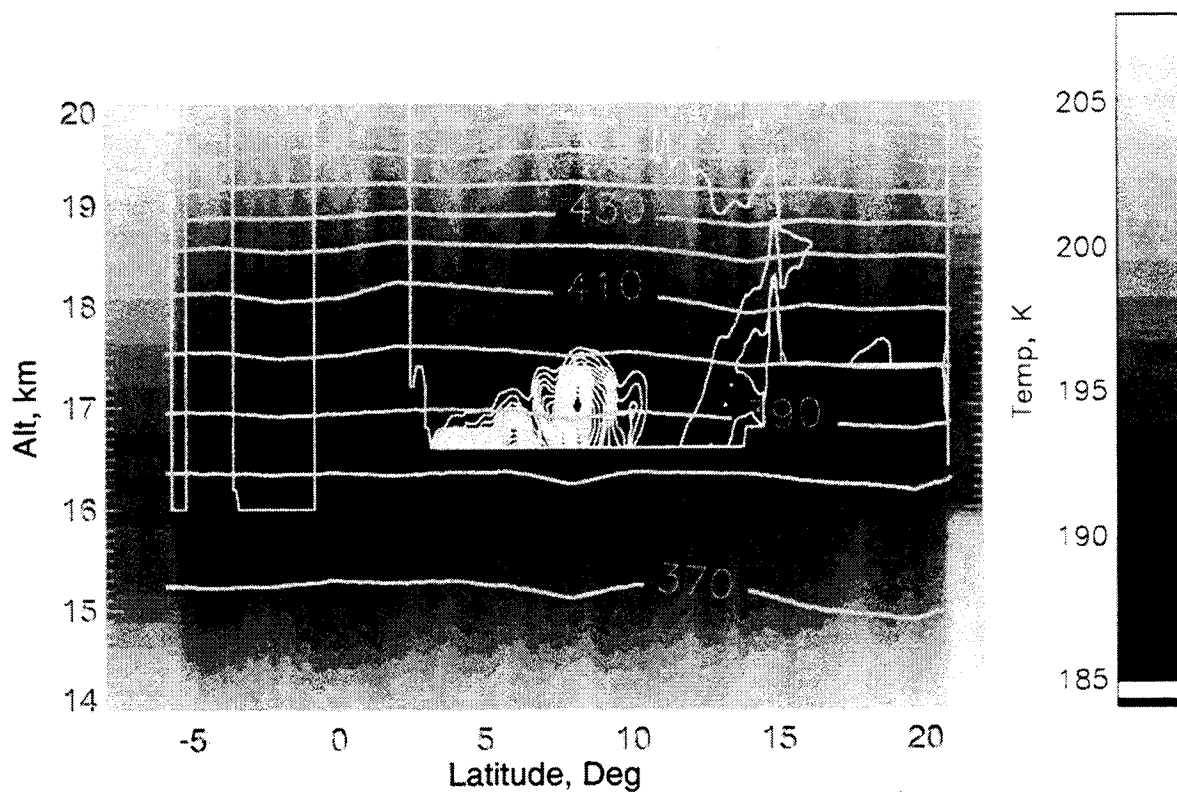


Fig 2c

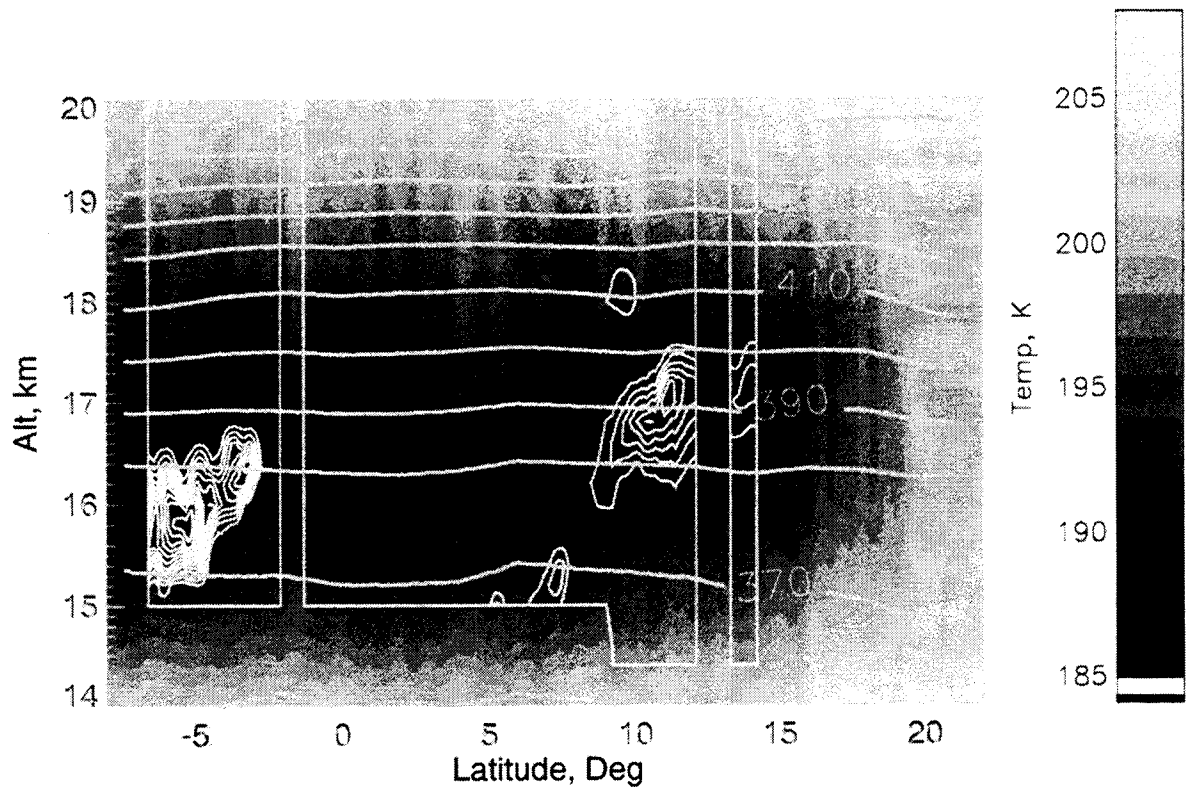


Fig 2d

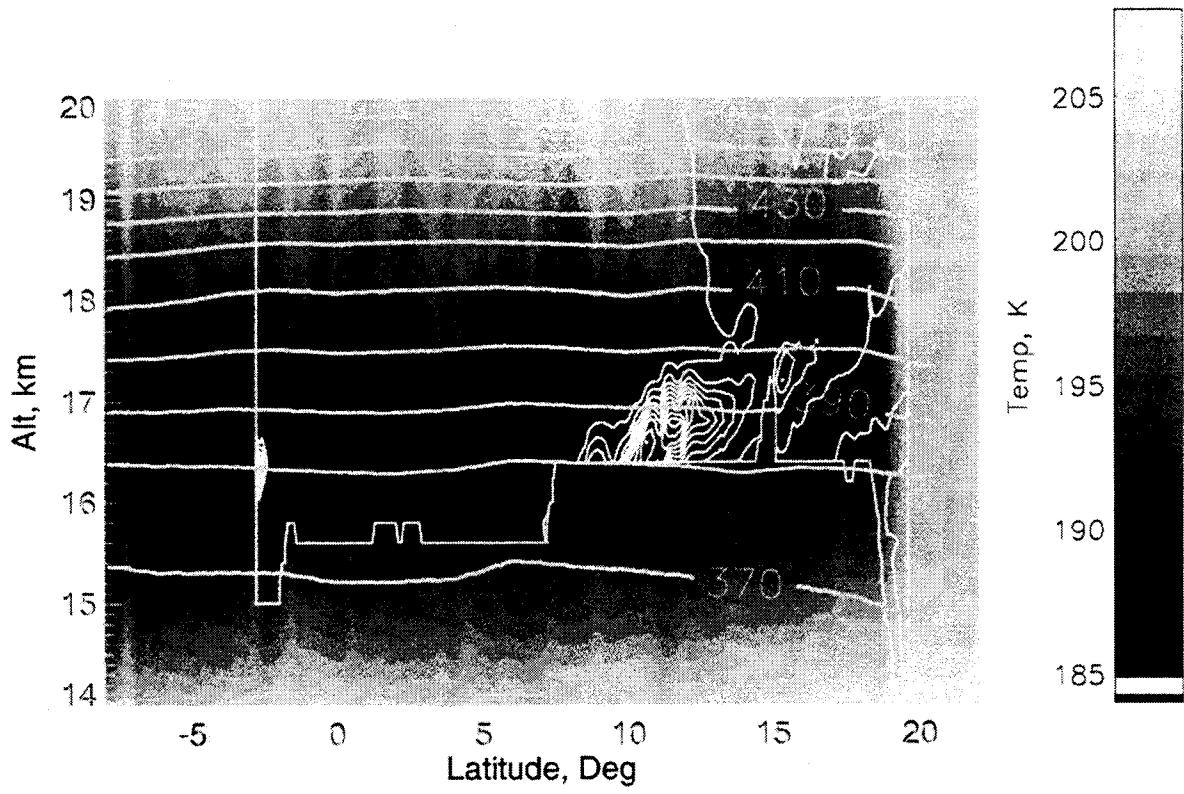


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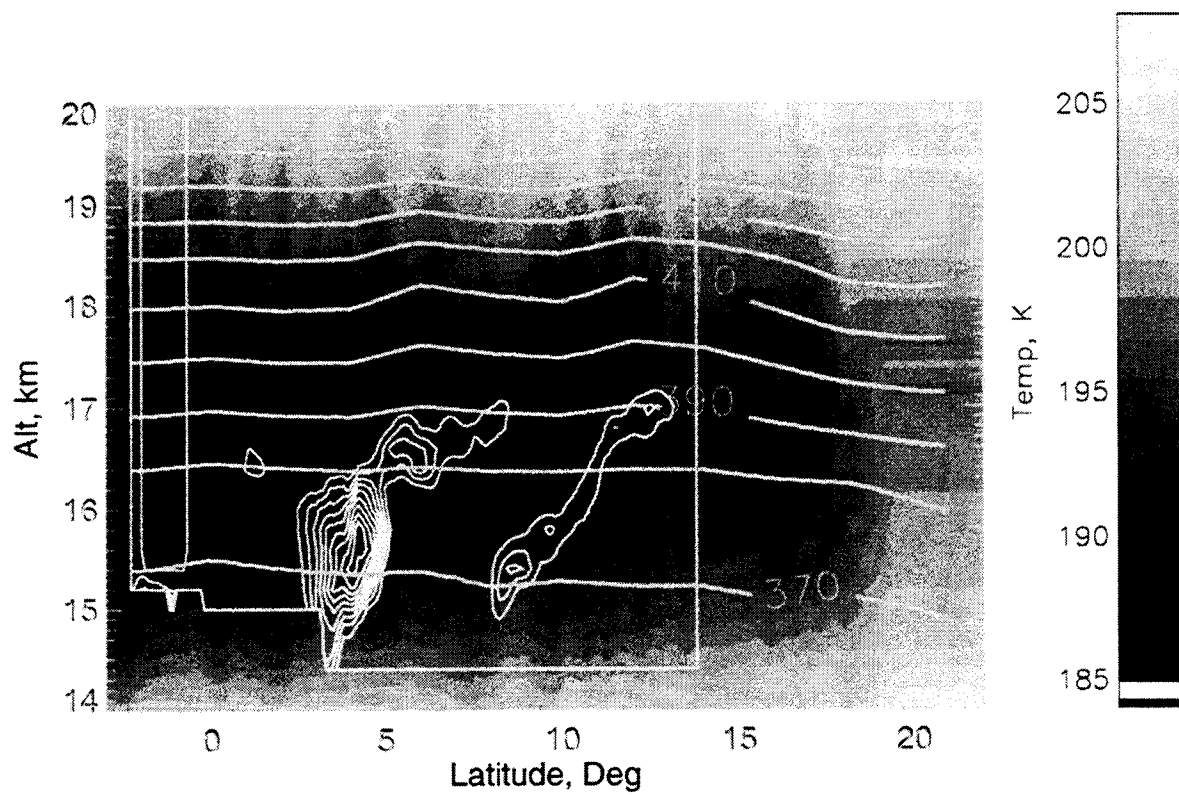


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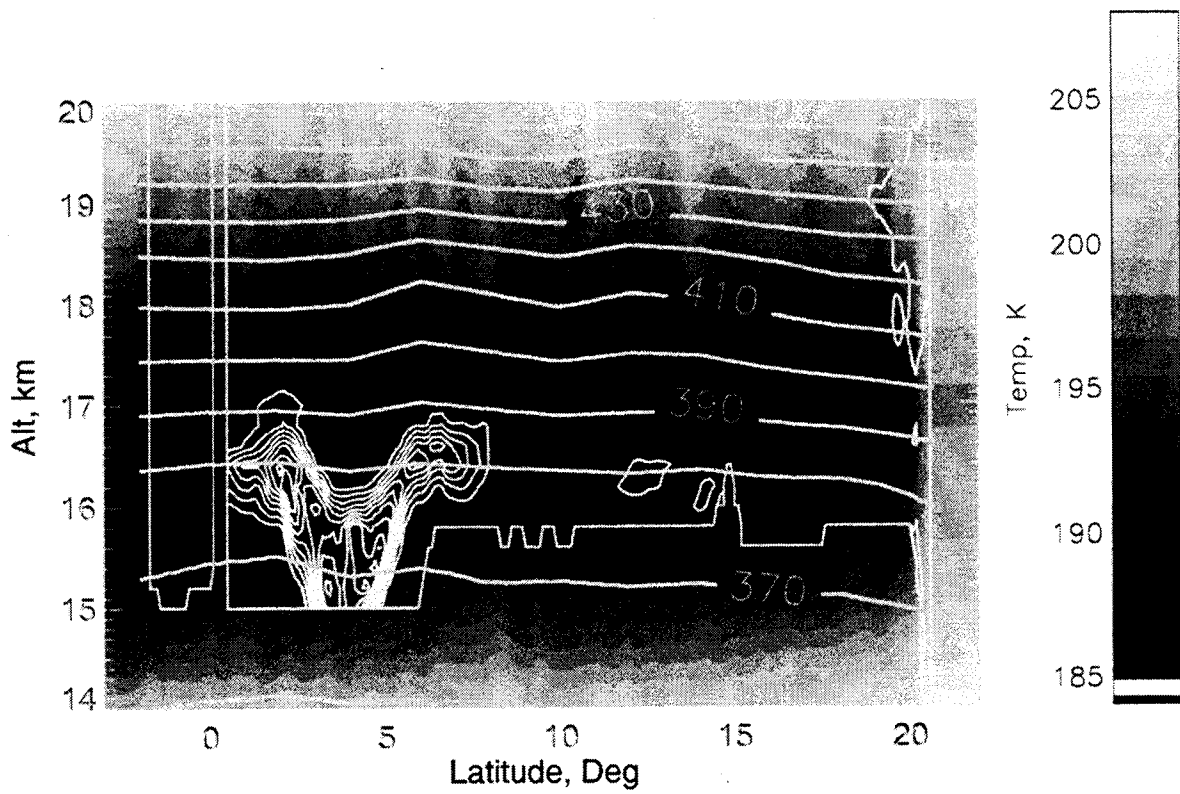


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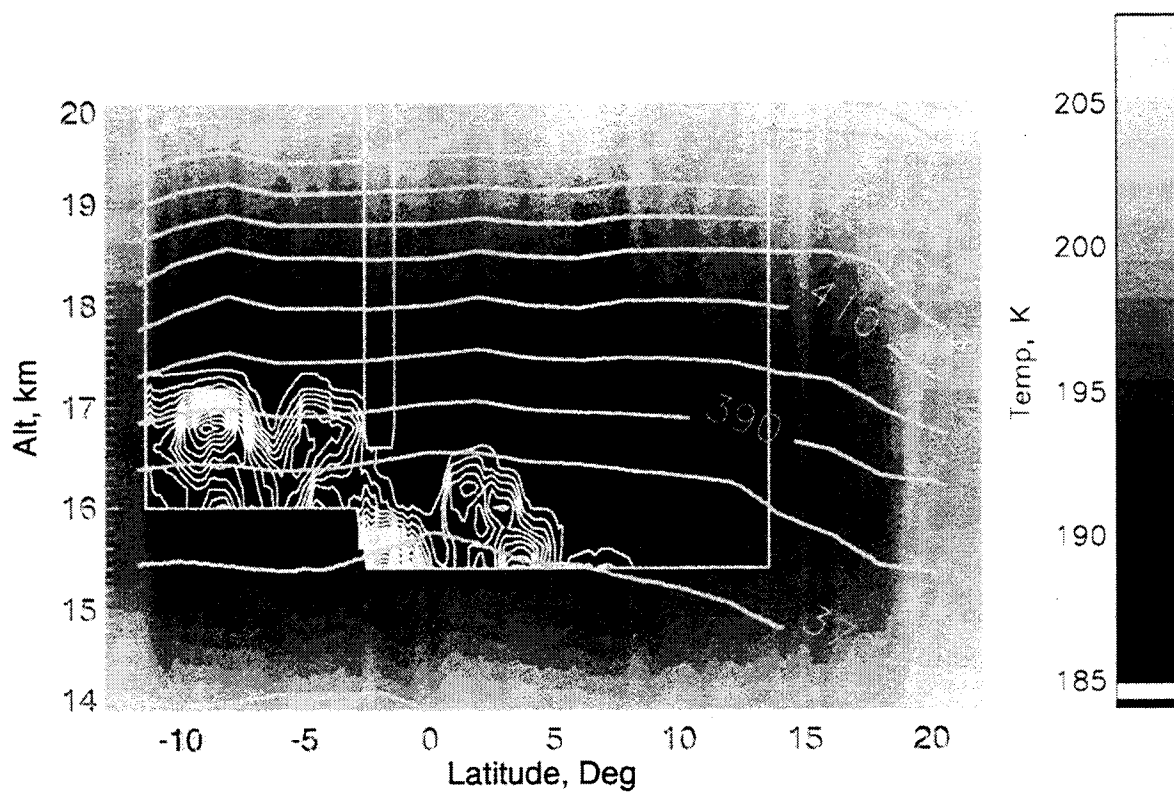
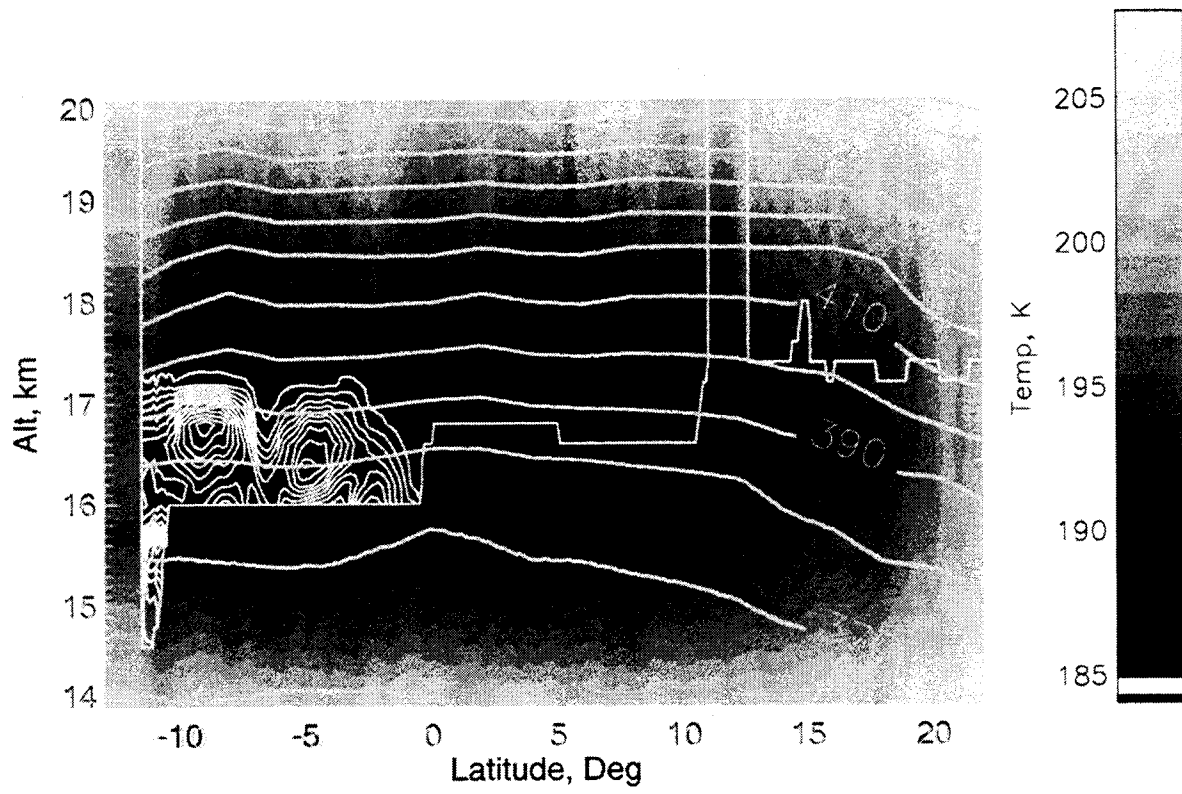


Fig 2h



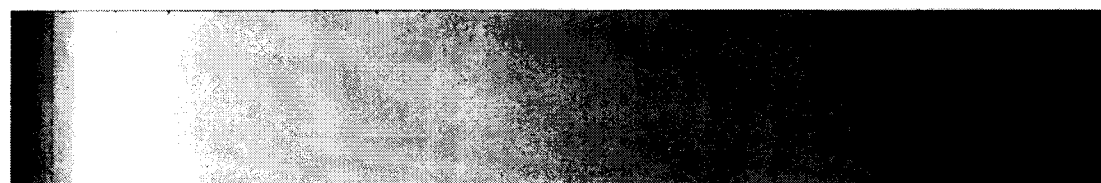
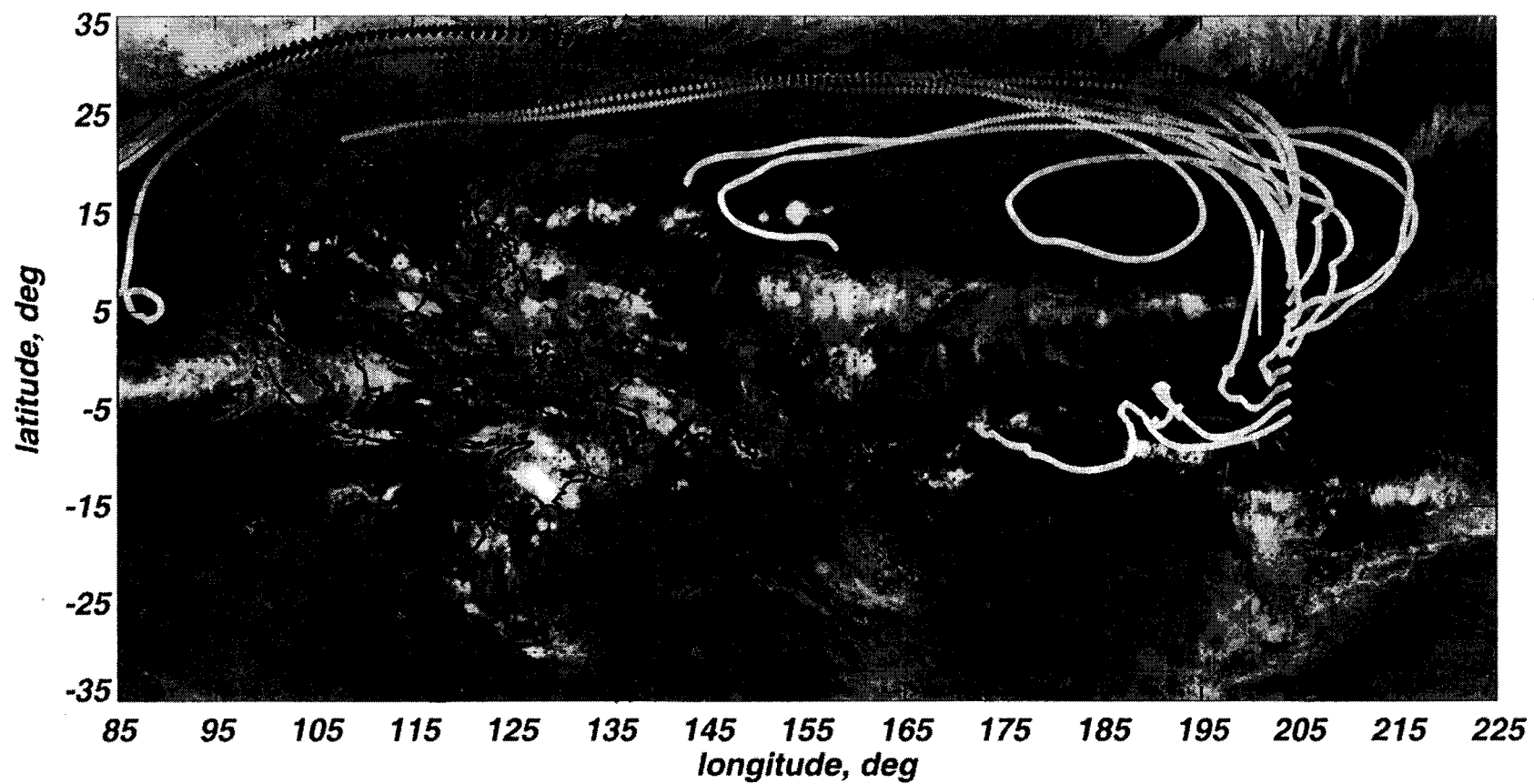
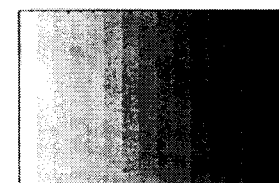


Image Temp



Traj Temp K

Fig. 3a

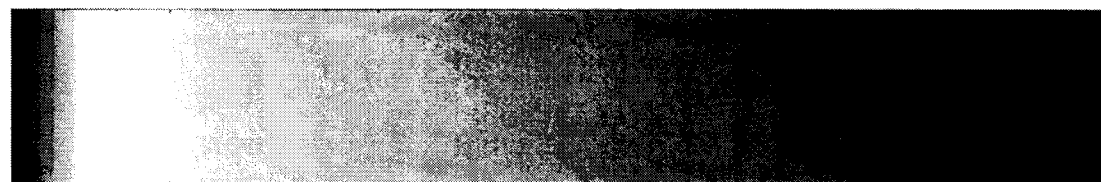
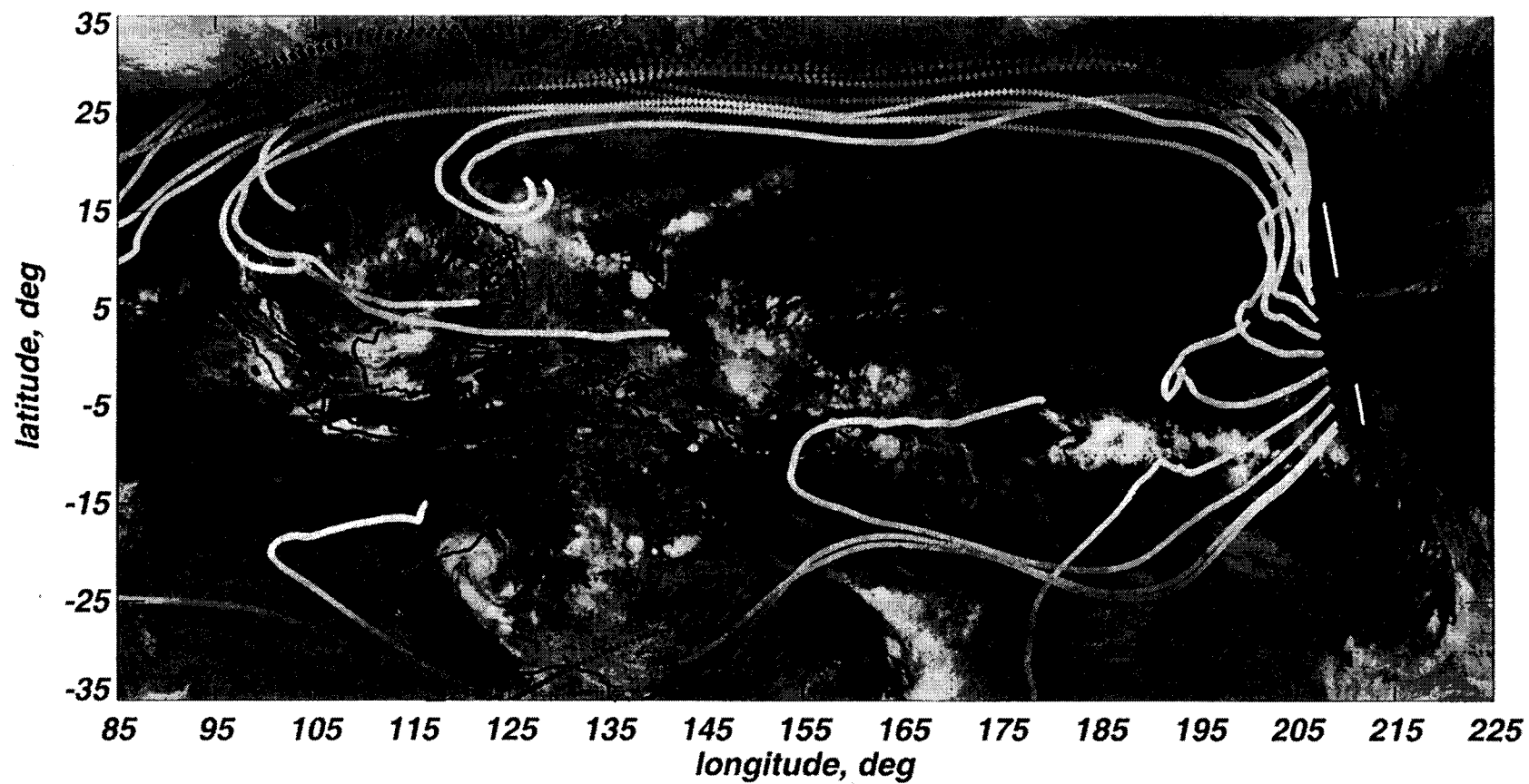
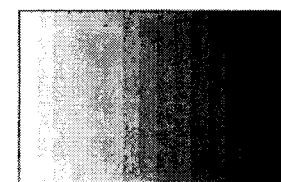


Image Temp



Traj Temp K

Fig. 3b

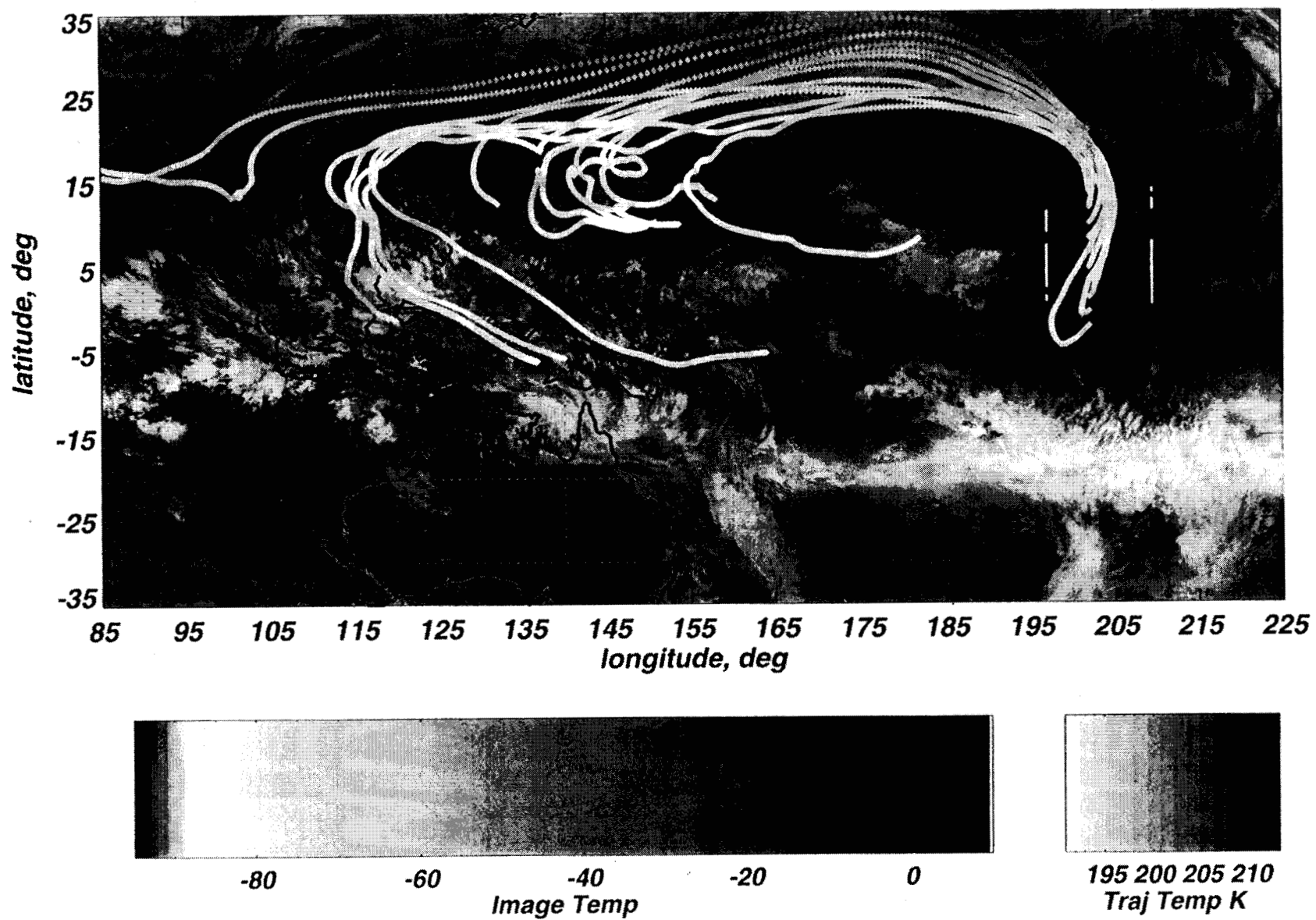


Fig. 3c

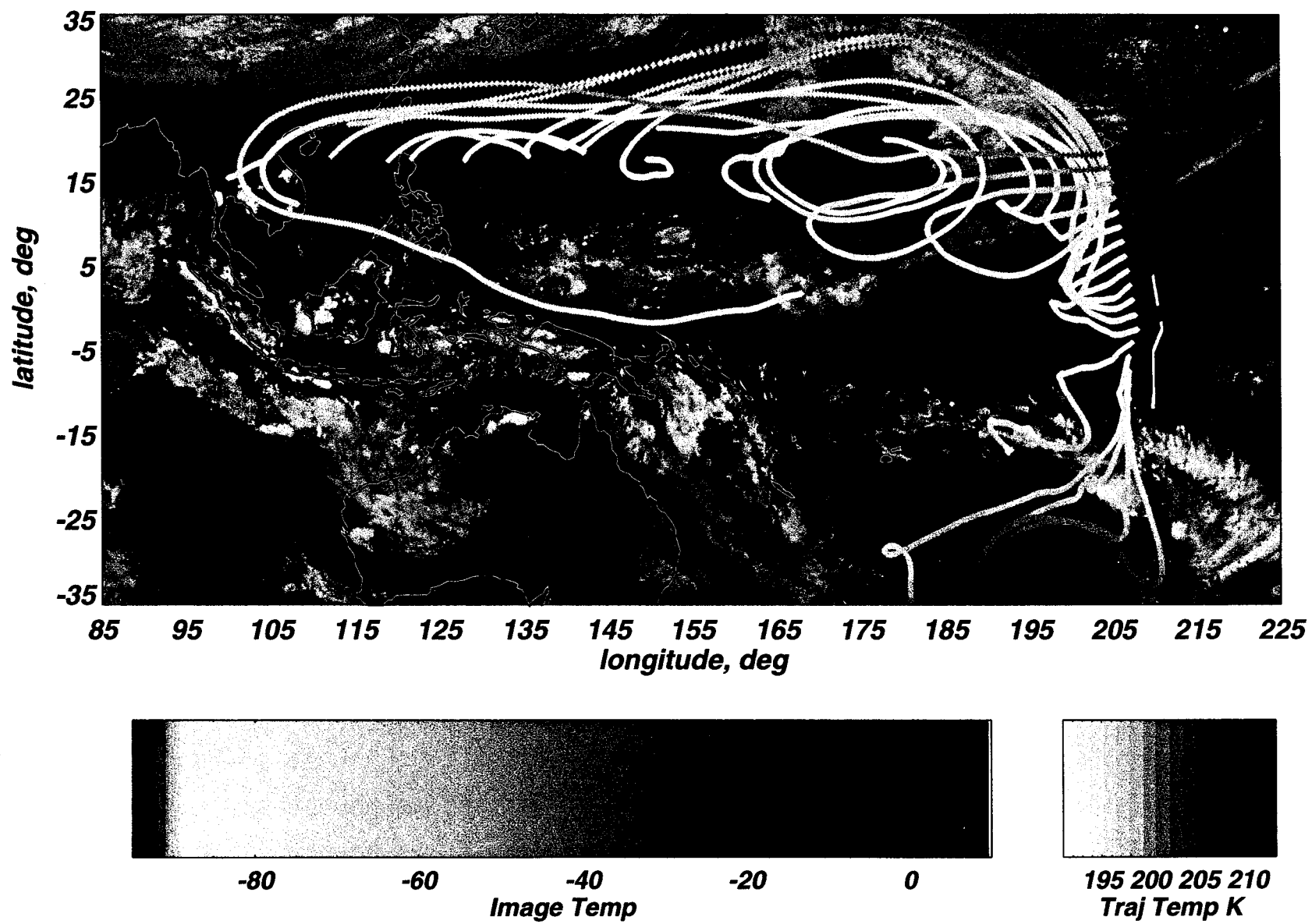


Fig. 3d

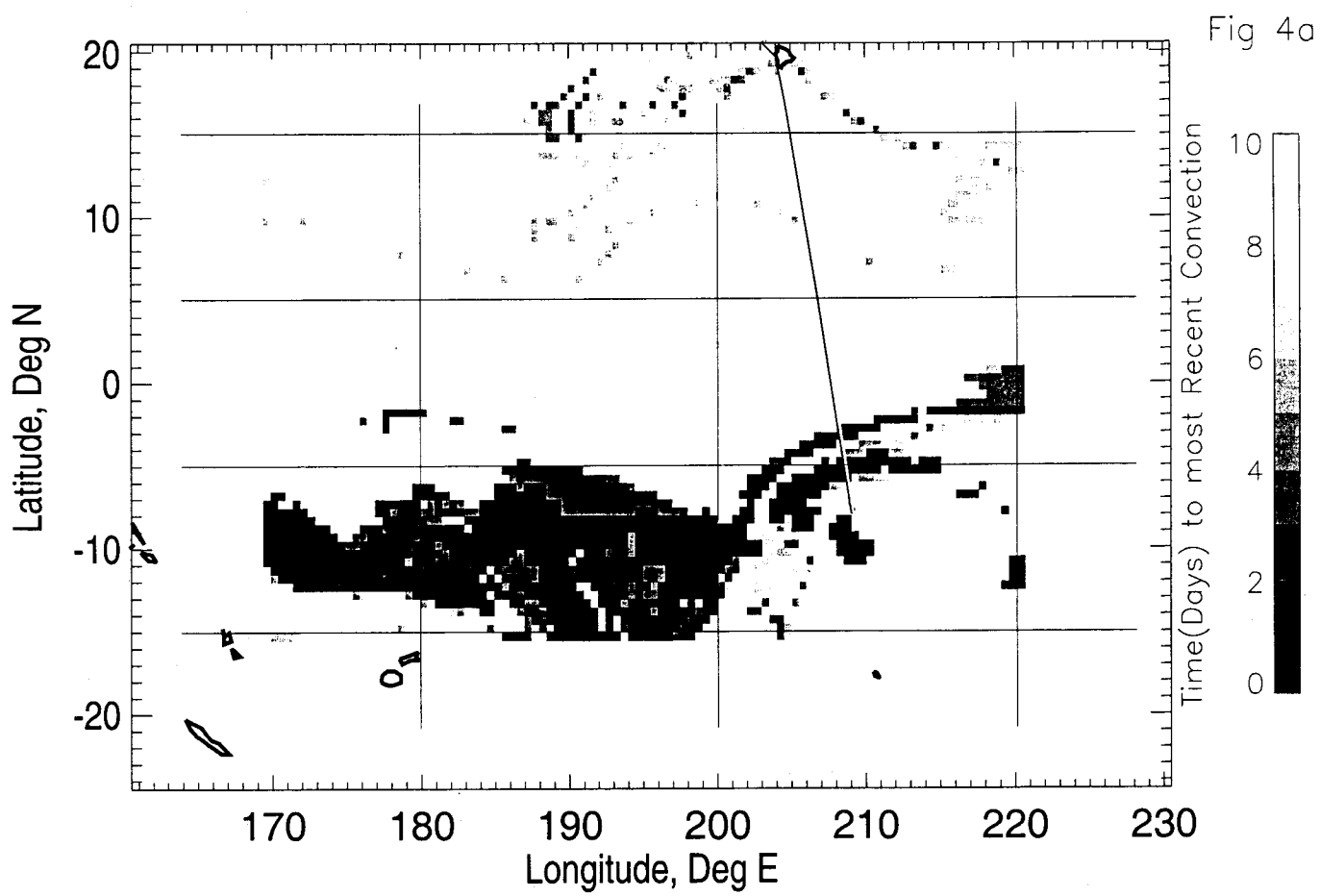


Fig 4b

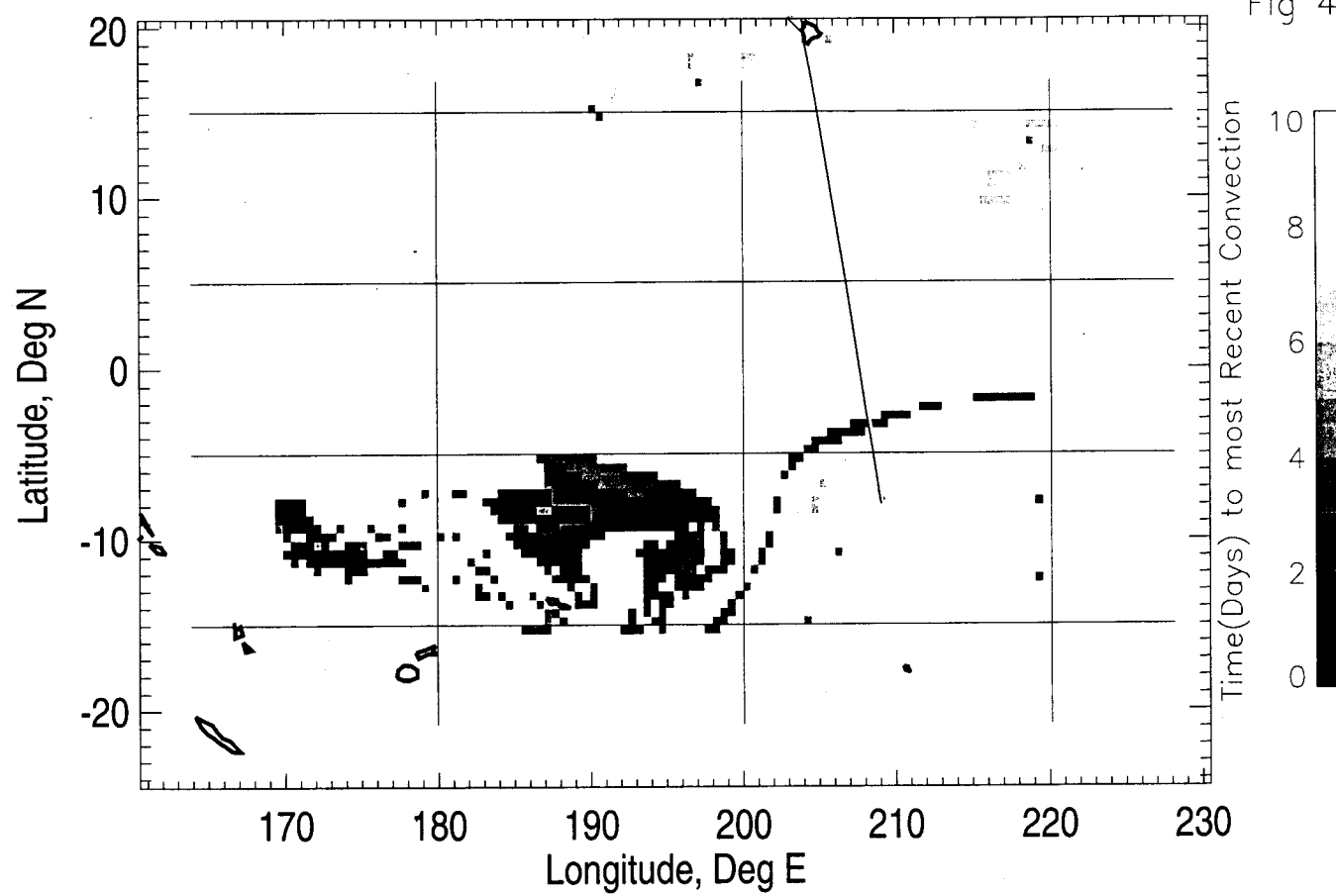


Fig 5a

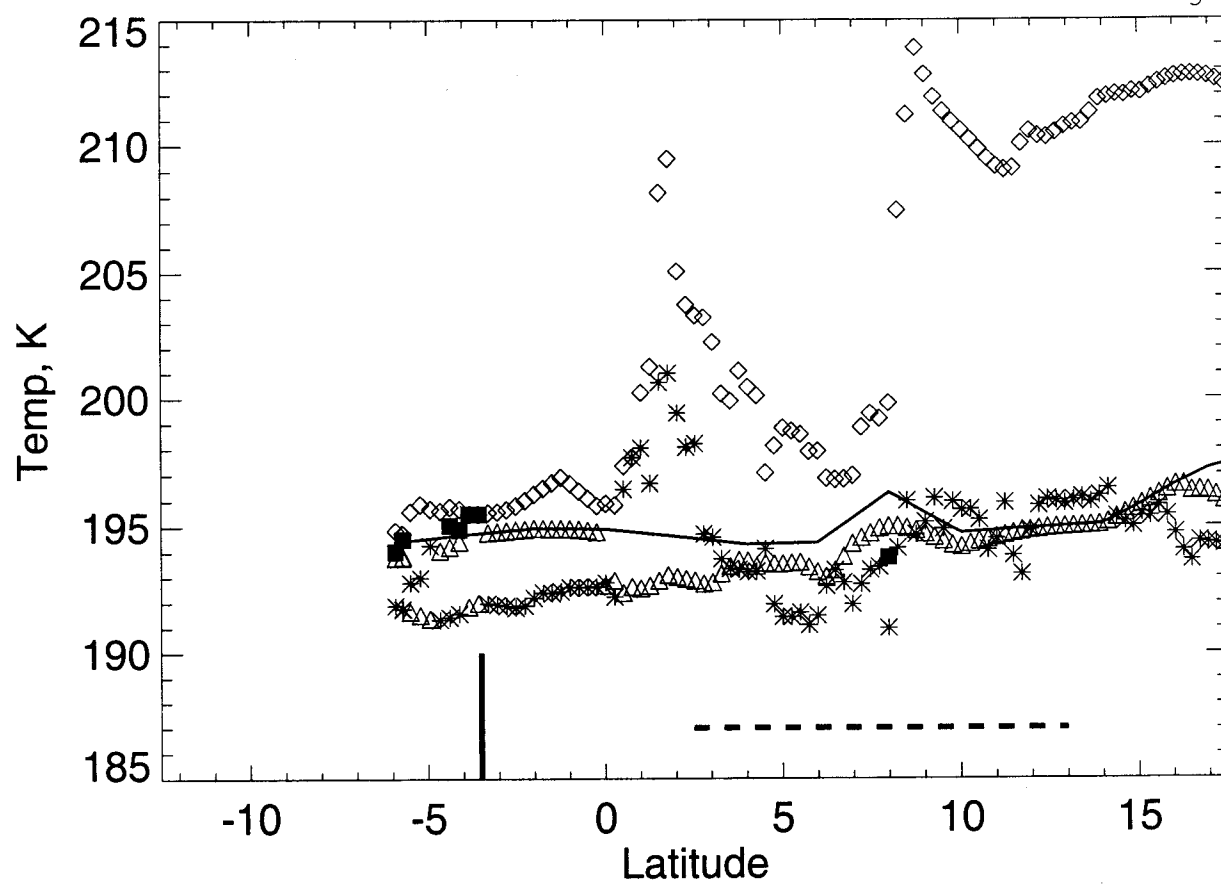


Fig 5b

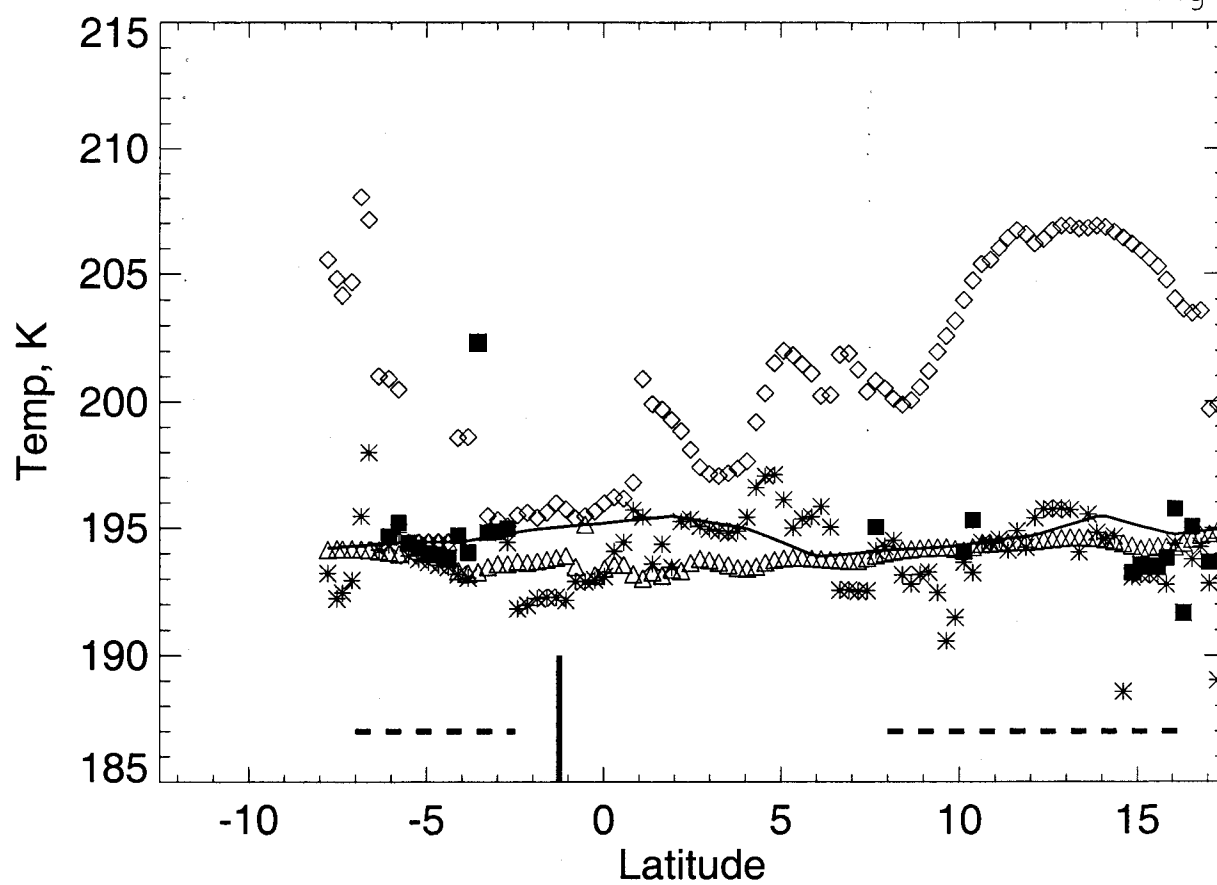


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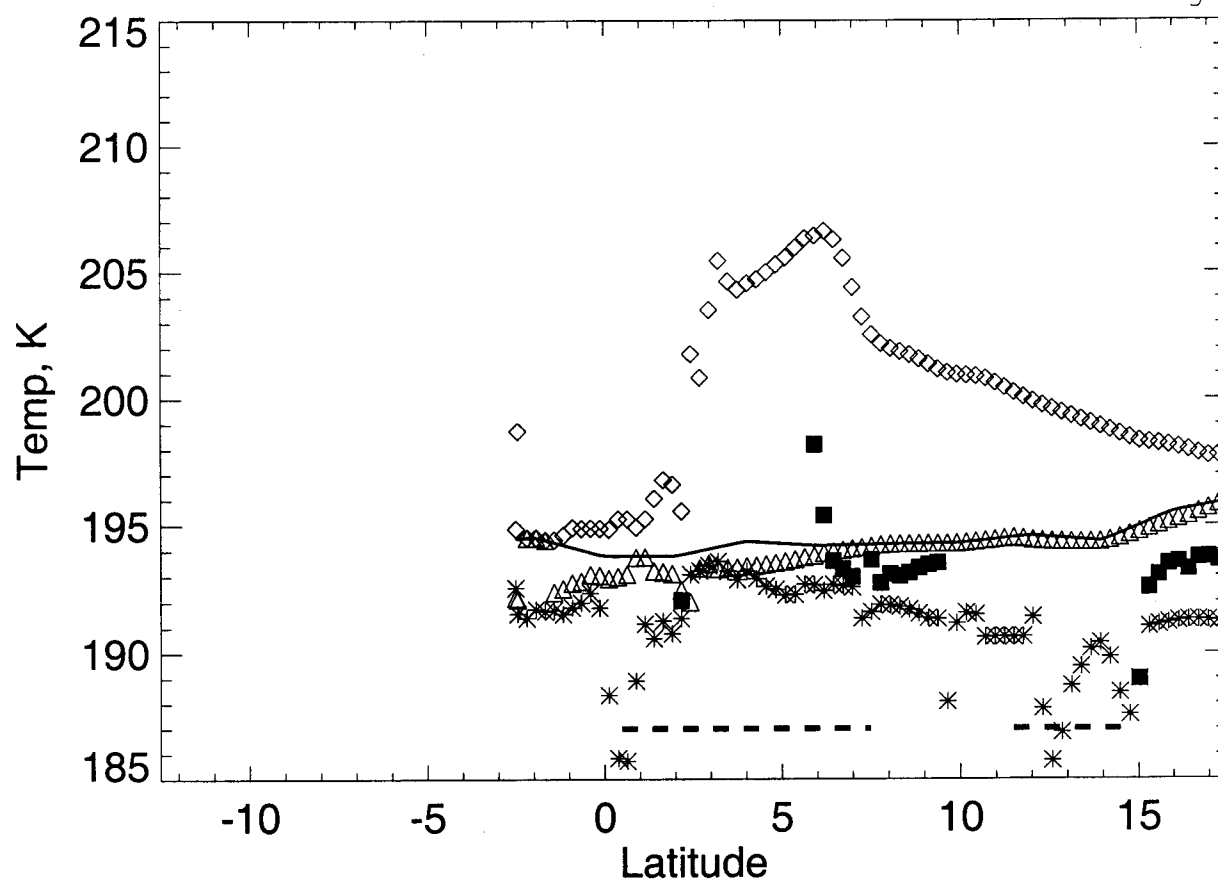


Fig 5d

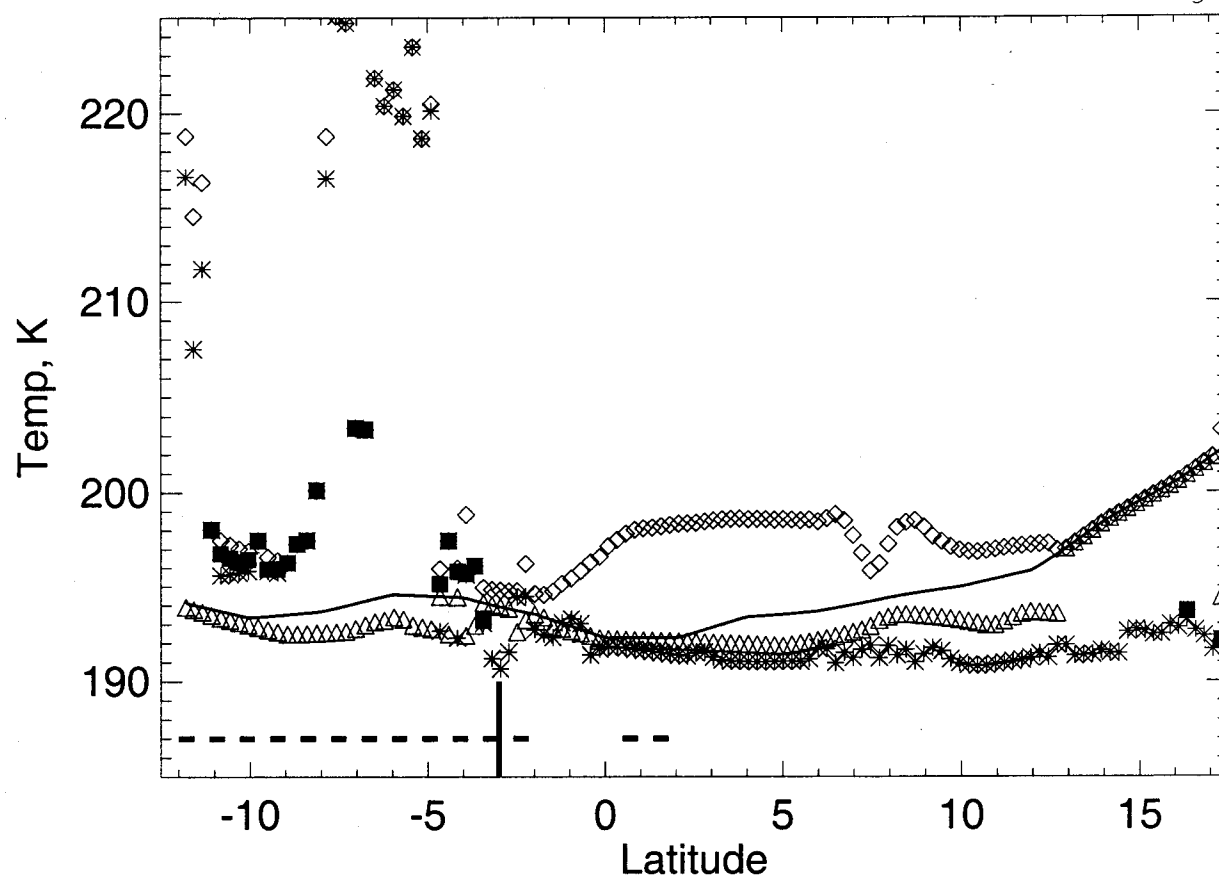


Fig 6

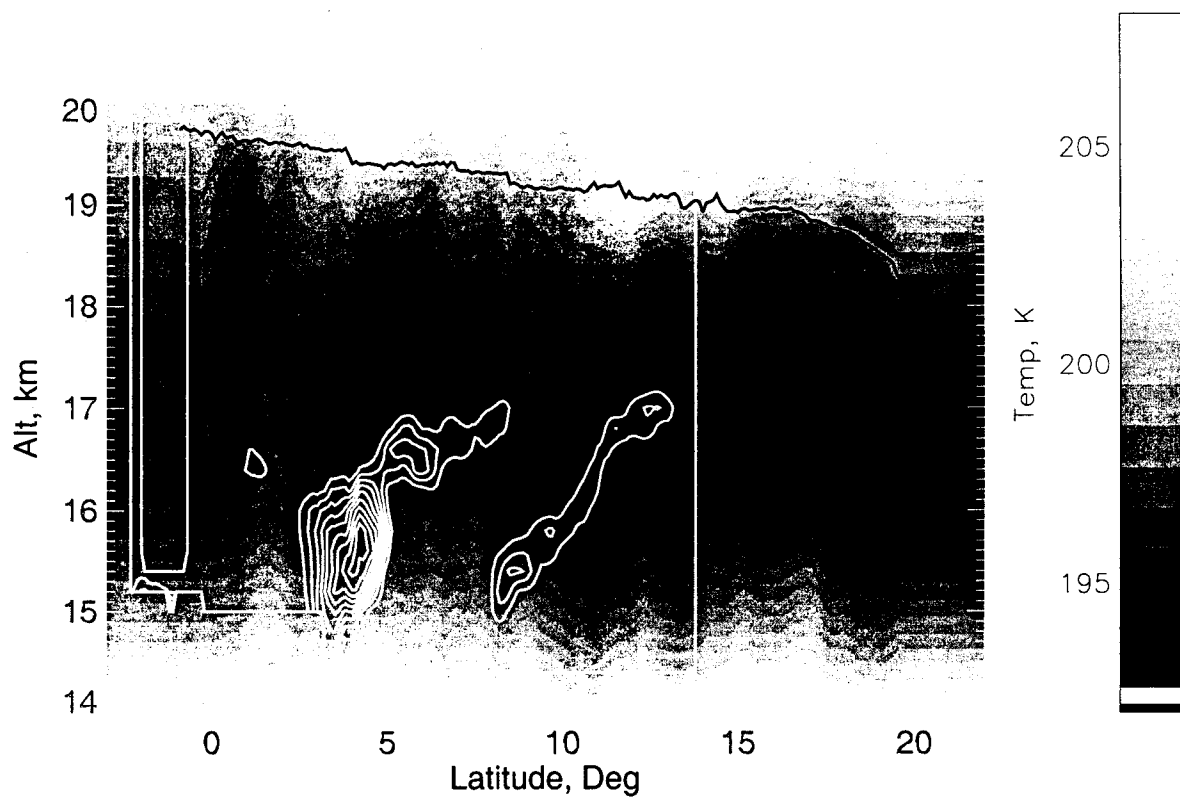


Fig 7

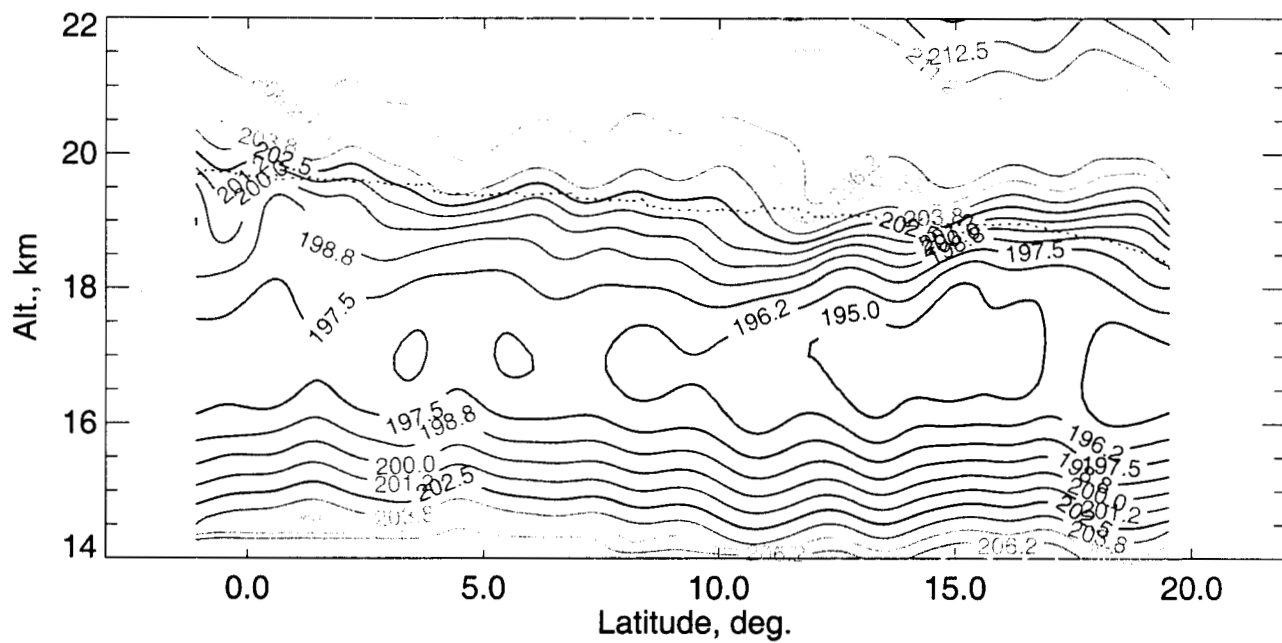
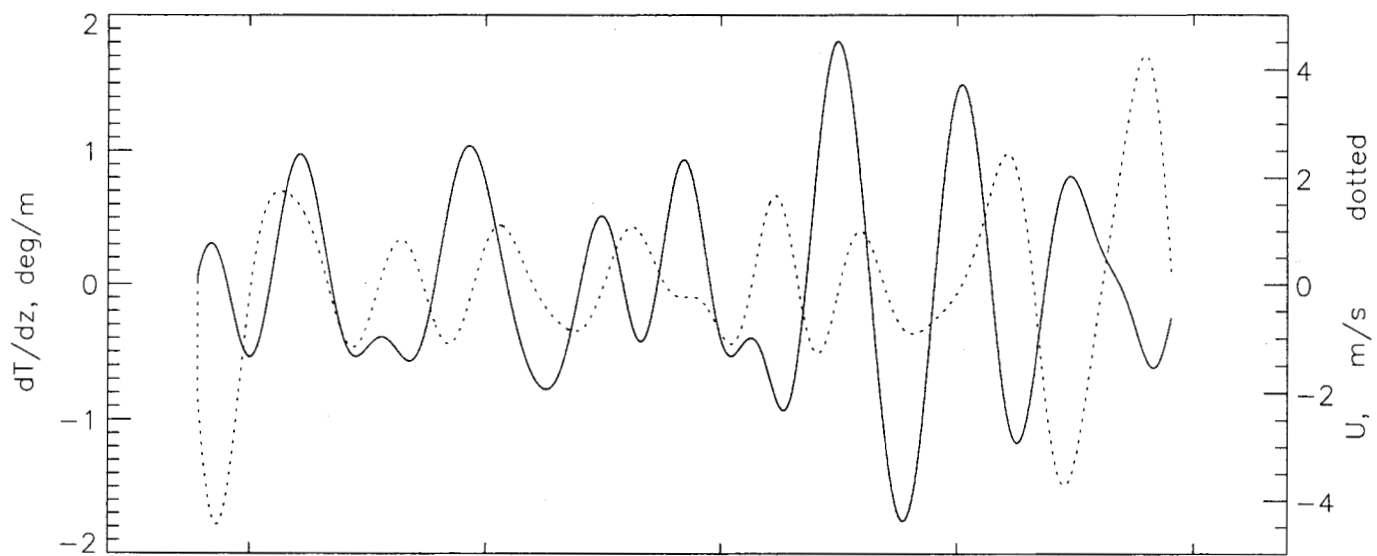
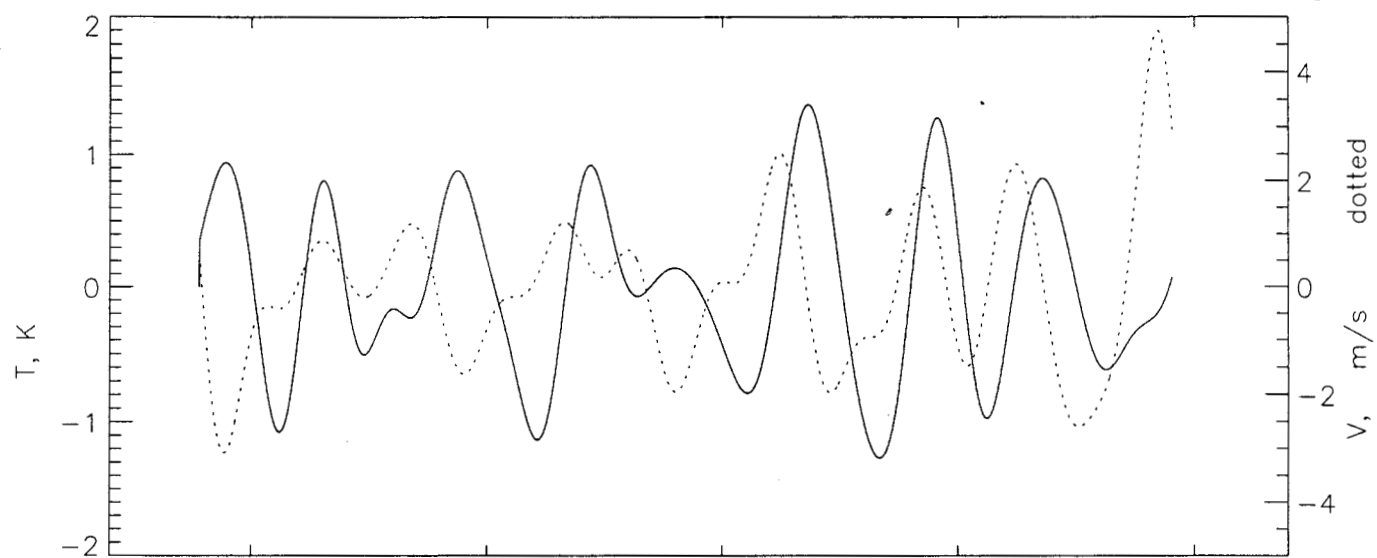


Fig 8

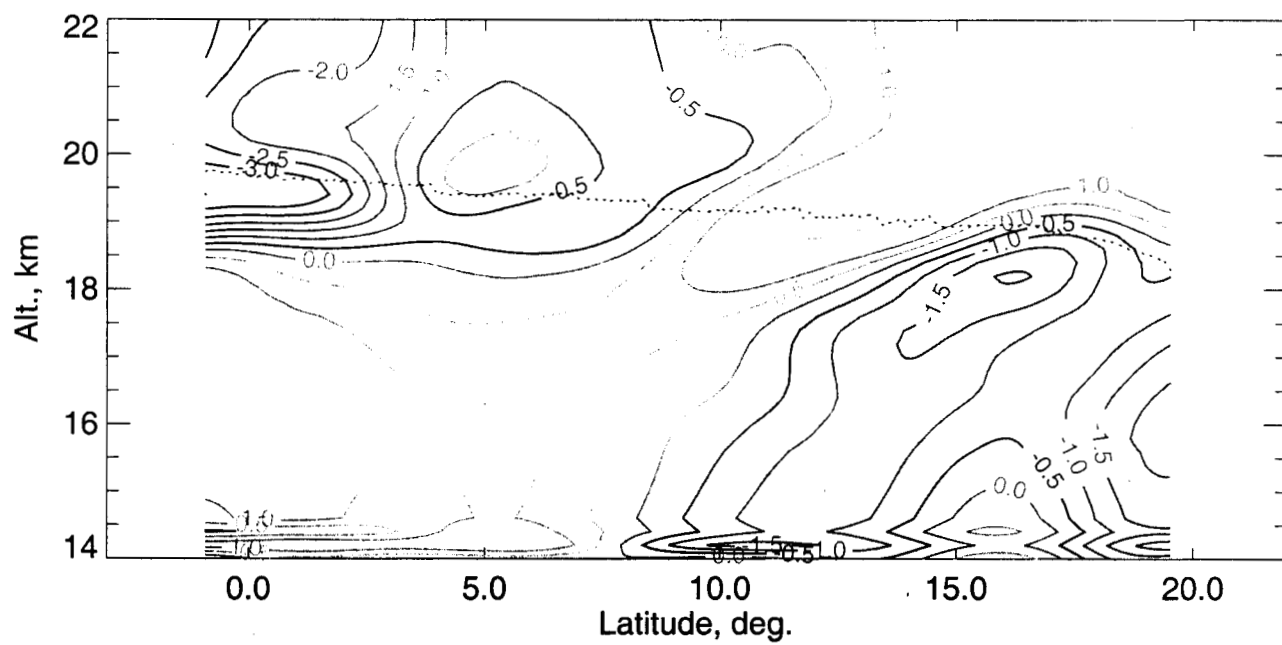
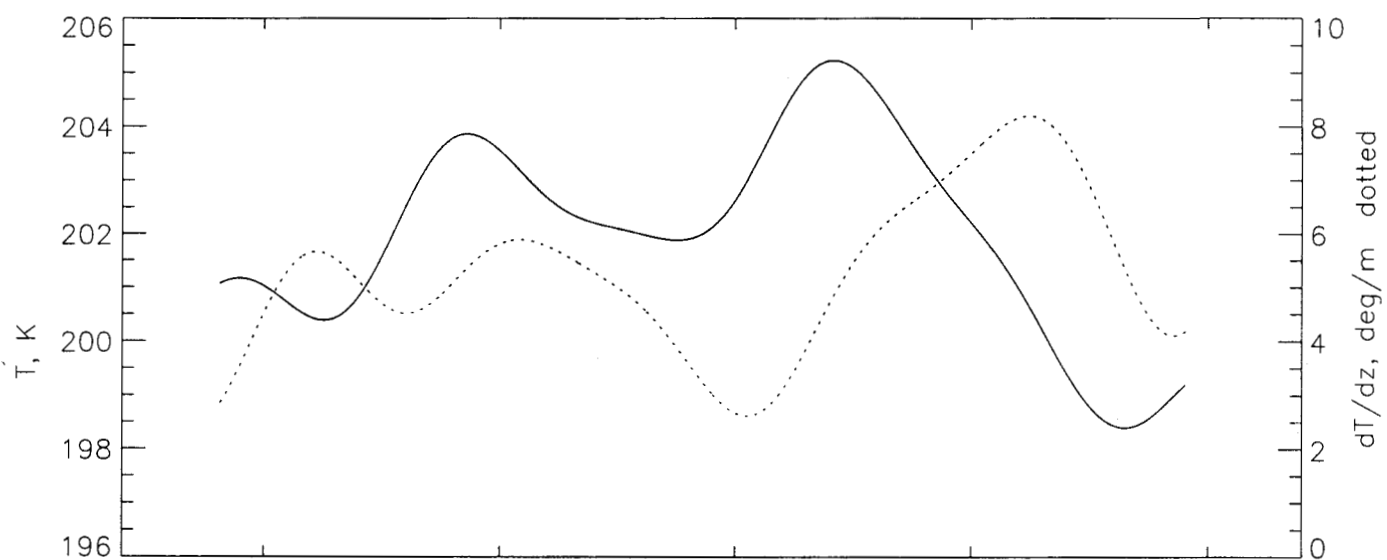
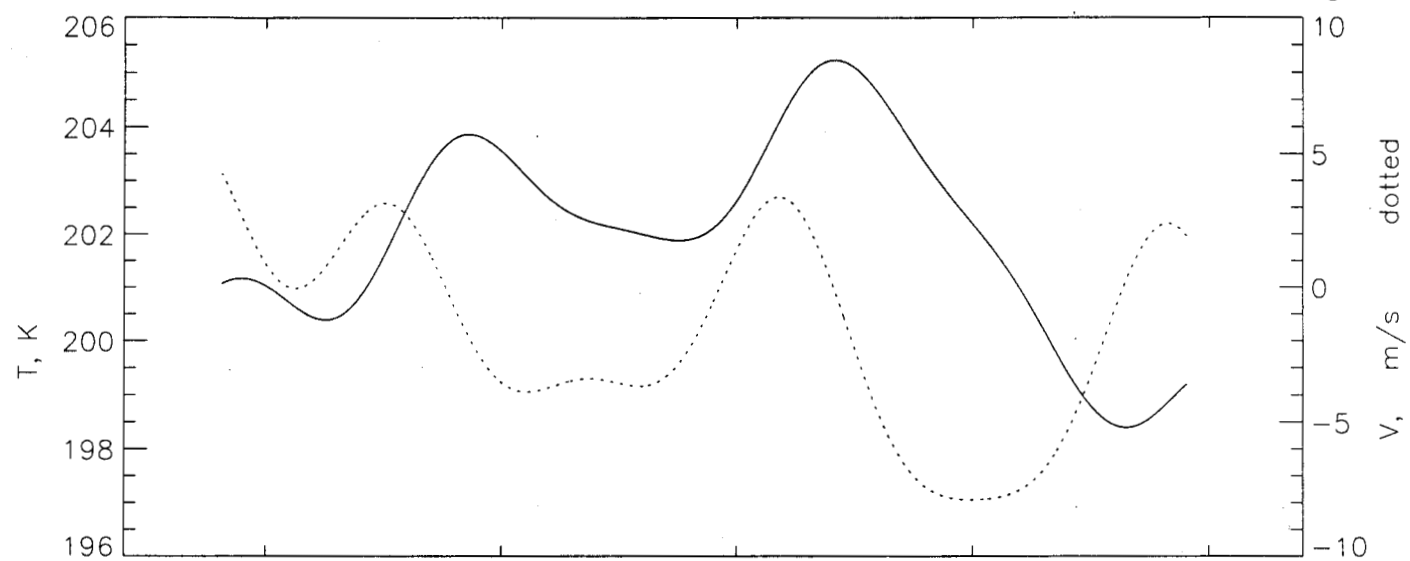


Fig 9a

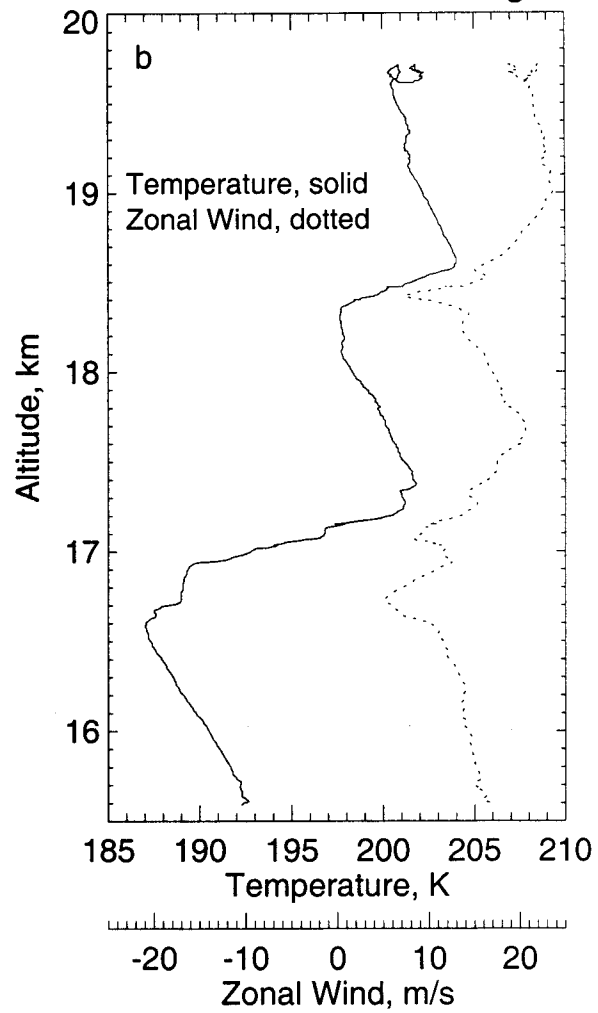
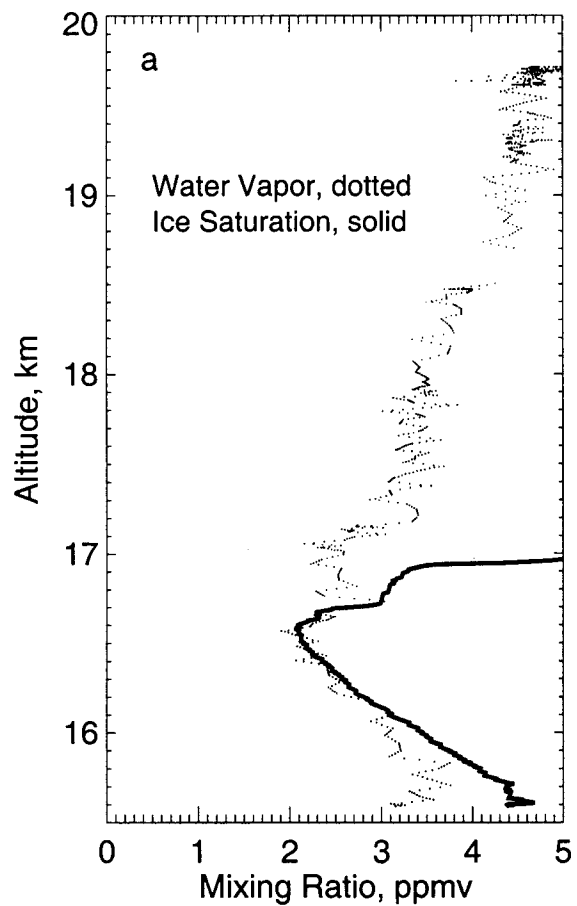


Fig 9b

